

Cosmology from quantum fluctuations of space-time (CMaDE model)

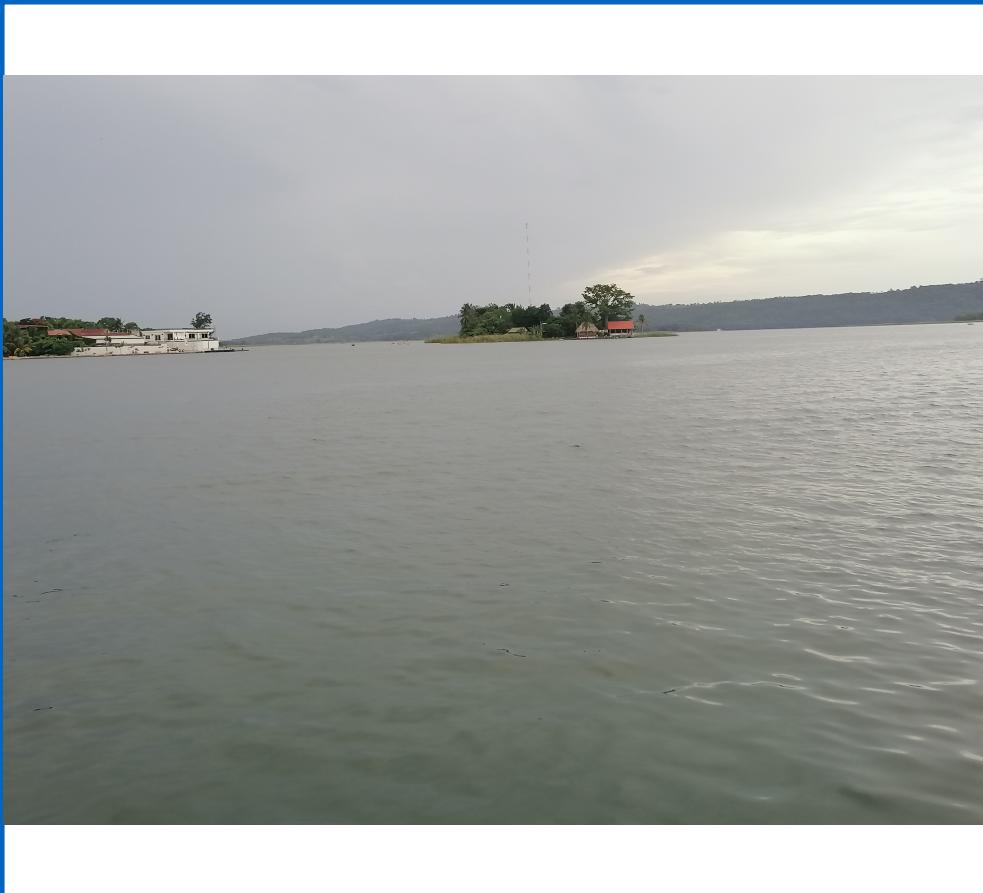


Tonatiuh Matos

<http://www.fis.cinvestav.mx/~tmatos/>

The CMaDE model

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The Compton mass

Tonatiuh Matos and Laura L-Parrilla. Rev. Mex.Fis., (2021), 67, 040703. arXiv:2108.05206

Energy of a photon of frequency ν

$$E = h\nu = mc^2$$

Compton wavelength

$$\lambda = \frac{h}{mc}$$

Graviton

$$m_g = 0$$

$$E = h\nu = mc^2$$



$$\lambda = \frac{h}{mc}$$

Everything gravitates

Compton mass

Tonatiuh Matos and Laura L-Parrilla. Rev. Mex.Fis., (2021), 67, 040703. arXiv:2108.05206

Weak field

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$



$$R_{\mu\nu} \sim \frac{1}{2} \square h_{\mu\nu} = \frac{1}{2} \square g_{\mu\nu}$$

Scalar Field

$$\square \Phi = \frac{m^2 c^2}{\hbar^2} \Phi = 0$$

Electromagnetic Field

$$\square A_\mu = \frac{m^2 c^2}{\hbar^2} A_\mu = 0$$

Gravitational Field

$$\square g_{\mu\nu} = \frac{m^2 c^2}{\hbar^2} g_{\mu\nu} = 0$$

Einstein Equations

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$$R_{\mu\nu} - \frac{1}{2}\mathcal{M}g_{\mu\nu}R = \kappa^2\mathcal{T}_{\mu\nu} = \kappa^2g_{\mu\nu}T$$

Einstein Equations

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$$R_{\mu\nu} - \mathcal{M}g_{\mu\nu} = 0$$

Weak field $\square g_{\mu\nu} - \frac{m^2 c^2}{\hbar^2} g_{\mu\nu} = 0$ $R_{\mu\nu} \sim \frac{1}{2} \square h_{\mu\nu} = \frac{1}{2} \square g_{\mu\nu}$

$$2R_{\mu\nu} - 2\mathcal{M}g_{\mu\nu} = 0$$

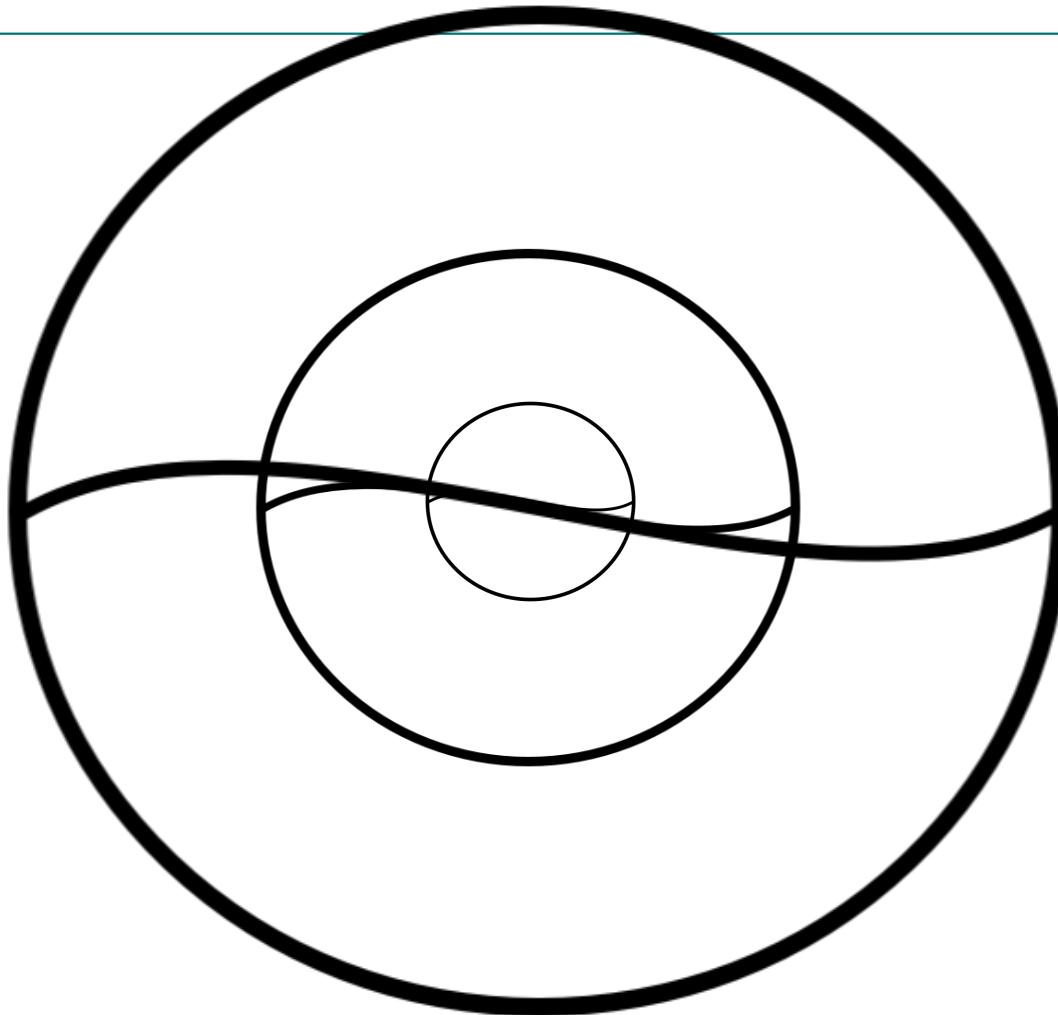
$$2\mathcal{M} = \frac{m^2 c^2}{\hbar^2} = \frac{4\pi^2}{\lambda^2}$$



Compton mass

$$m = \frac{h}{\lambda c}$$

The Universe is finite



$$ds^2 = -c^2 dt^2 + ad\mathbf{x}^2 = 0$$

The CMaDE model

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$$R_H = H_0 \int_0^{today} \frac{dt}{a} = H_0 \int_{-\infty}^0 \frac{dN}{H} e^{-N}$$

$$N = \ln(a) \quad \dot{N} = H$$

$$\lambda = \frac{c}{H_0} R_H$$

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + \mathcal{M}g_{\mu\nu} = \kappa^2 T_{\mu\nu}$$

$$\mathcal{M} = \frac{2\pi^2}{\lambda^2} = \frac{2\pi^2}{R_H^2} \frac{H_0^2}{c^2}$$

The Freedman equation

T.M., Luis Escamilla, Alberto Vázquez and Maribel Hernández-Márquez. In preparation.

$$H^2 = \frac{\kappa^2}{3} (\rho_m + \rho_r + \rho_\Lambda)$$

The Freedman equation

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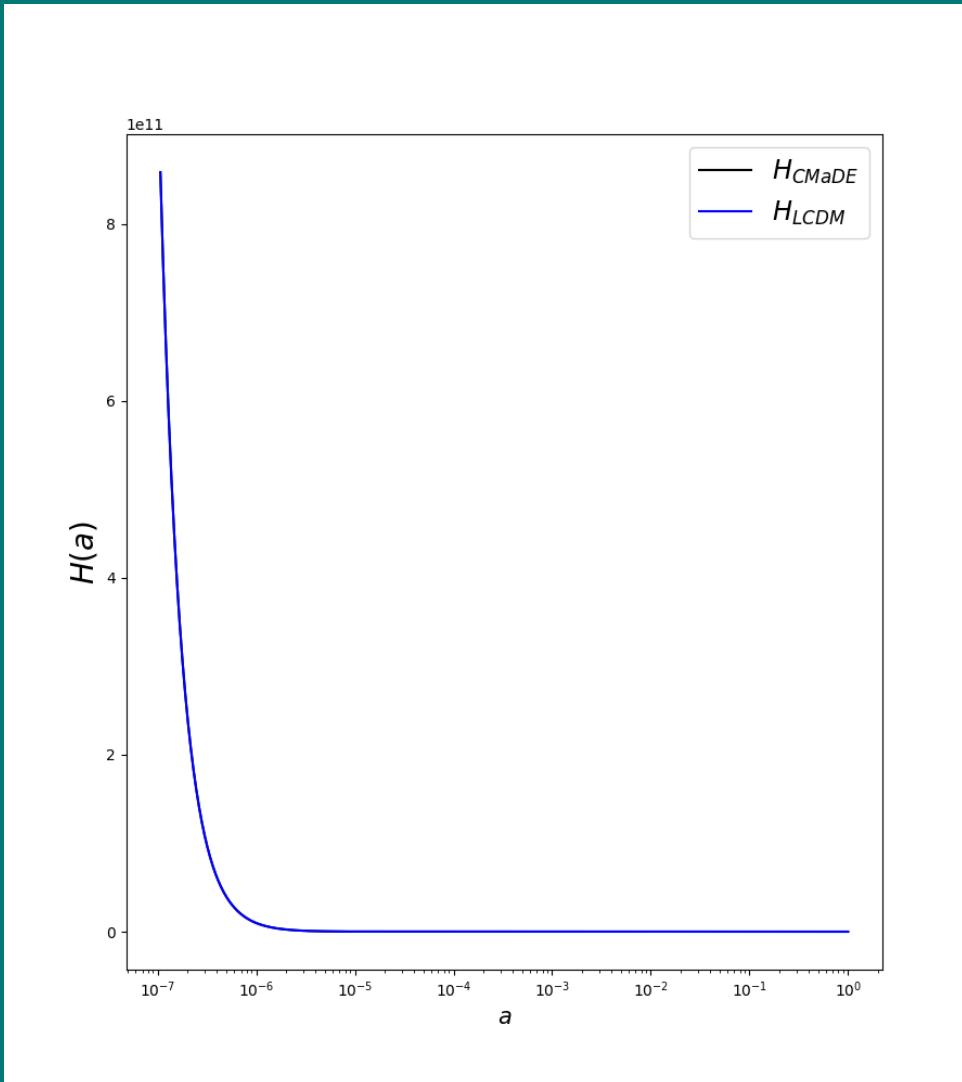
$$H^2 + \frac{k}{a^2} - \frac{\mathcal{M}c^2}{3} = \frac{\kappa^2 \rho}{3}$$

$$\Omega_{\mathcal{M}} = \frac{1}{3} \frac{c^2}{H_0^2} \frac{2\pi^2}{\lambda^2} = \frac{2}{3} \frac{\pi^2}{R_H^2}$$

$$\begin{aligned}
 \mathcal{H}^2 &= \Omega_b^0 + \Omega_{DM}^0 + \Omega_r^0 + \Omega_{0k}^0 e^{-2N} + \Omega_{\mathcal{M}}^0 \\
 \Omega_{DM}^{0'} &= -3\Omega_{DM}^0 - k_c \Omega_{\mathcal{M}}^{0'} \\
 \Omega_b^{0'} &= -3\Omega_b^0 \\
 \Omega_r^{0'} &= -4\Omega_r^0 \\
 \Omega_{\mathcal{M}}^{0'} &= Q \frac{\sqrt{6}}{\pi} \Omega_{\mathcal{M}}^0 {}^{3/2} \frac{e^{-N}}{\mathcal{H}}
 \end{aligned}$$

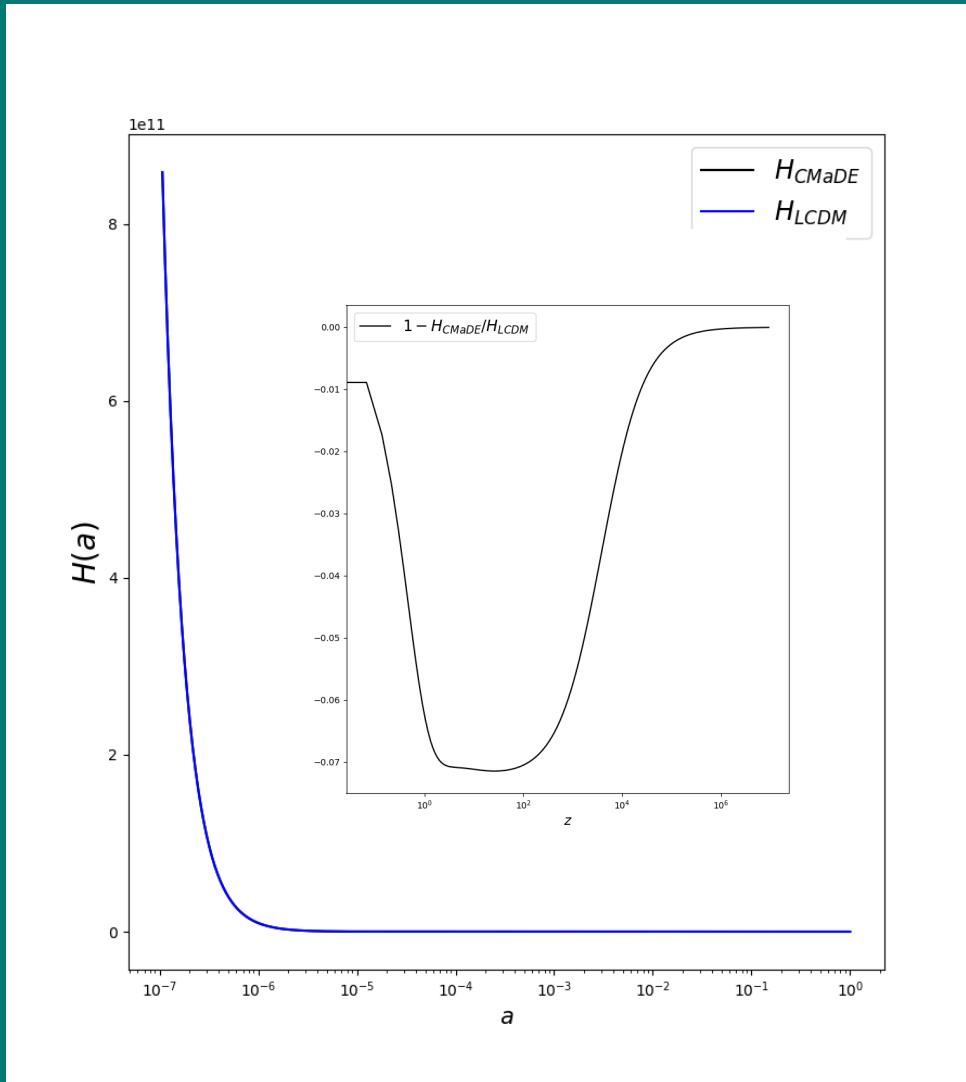
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$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \frac{2\pi^2}{\lambda^2}g_{\mu\nu} = \kappa^2 T_{\mu\nu} \quad \lambda = \text{The Compton wavelength of the graviton}$$

$$\mathcal{M} = \frac{2\pi^2}{\lambda^2}$$

The Friedman equations

$$\begin{aligned}\mathcal{H}^2 &= \Omega_b^0 + \Omega_{DM}^0 + \Omega_r^0 + \Omega_{0k}^0 e^{-2N} + \Omega_{\mathcal{M}}^0 \\ \Omega_{DM}^{0'} &= -3\Omega_{DM}^0 - k_c \Omega_{\mathcal{M}}^{0'} \\ \Omega_b^{0'} &= -3\Omega_b^0 \\ \Omega_r^{0'} &= -4\Omega_r^0 \\ \Omega_{\mathcal{M}}^{0'} &= Q \frac{\sqrt{6}}{\pi} \Omega_{\mathcal{M}}^0 {}^{3/2} \frac{e^{-N}}{\mathcal{H}}\end{aligned}$$

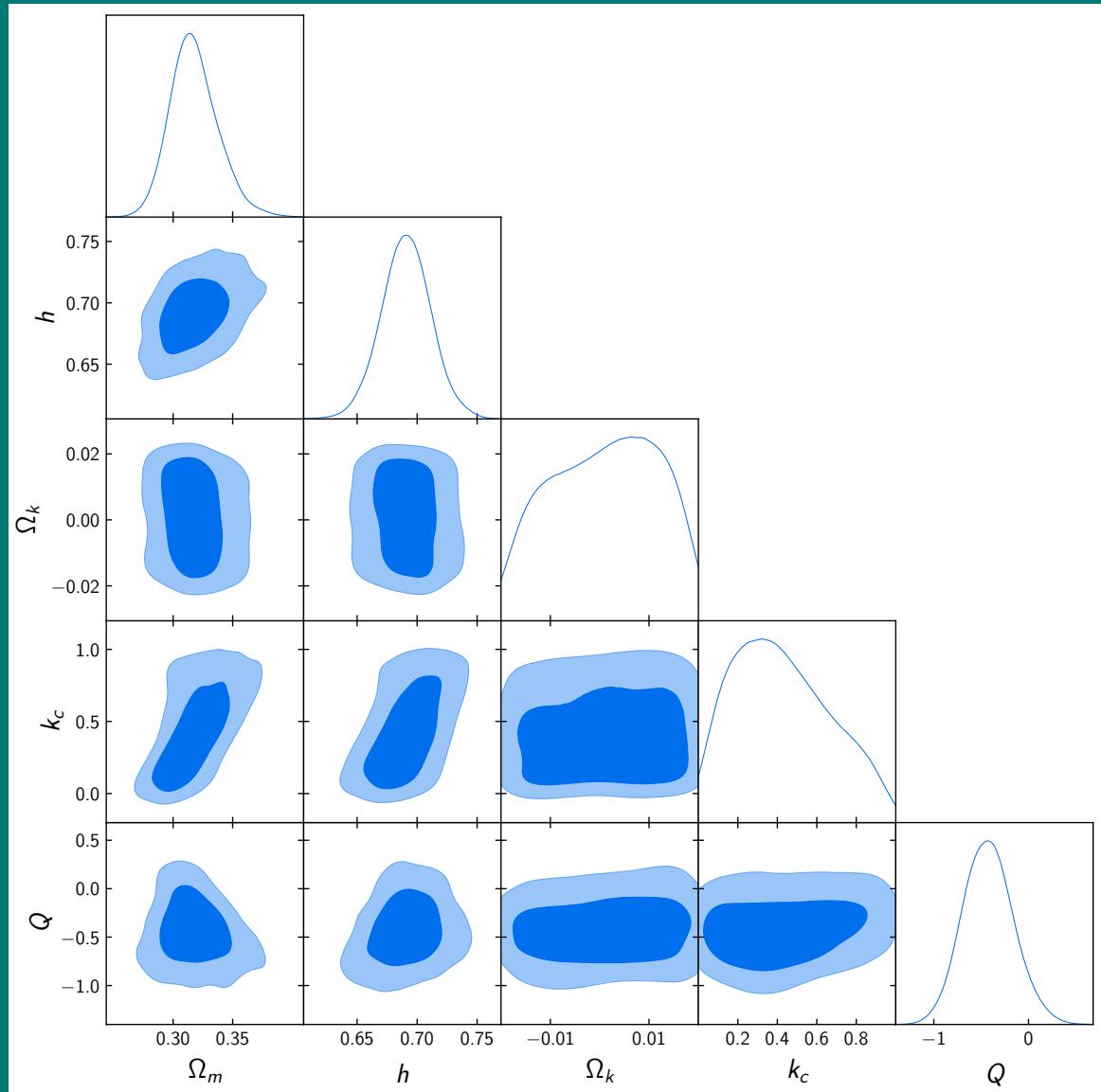
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CC+BAO+ Pantheon

$$\Delta\chi^2 = 3.25$$

in favor of CMaDE against Λ CDM

$$h = 0.69, \Omega_m = 0.31, \Omega_k = 0.08$$

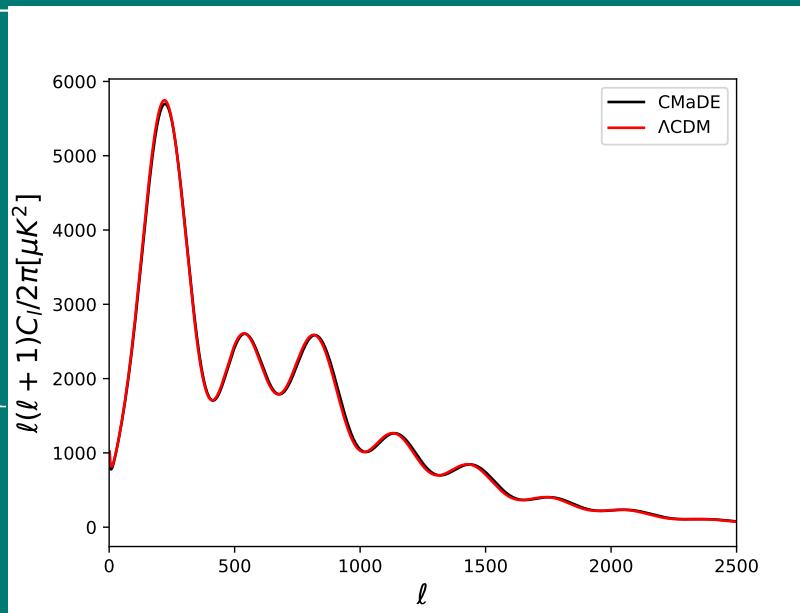


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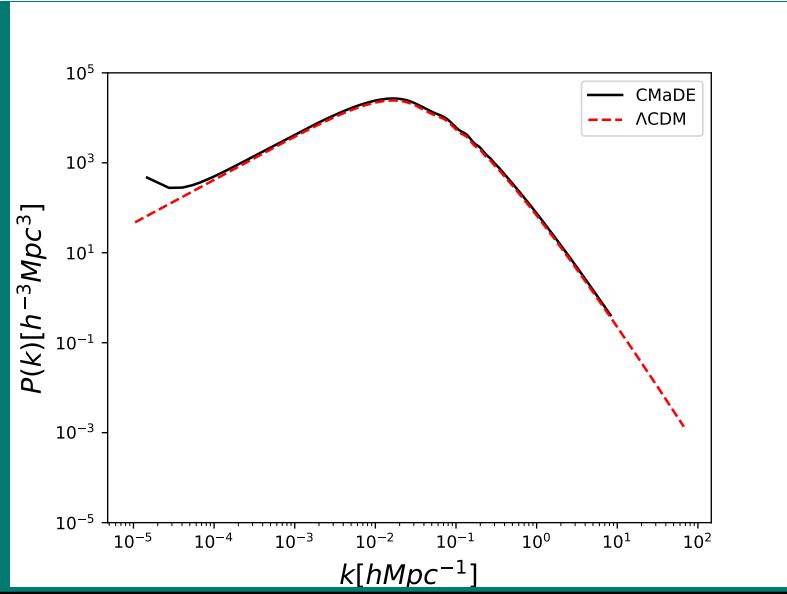
CMB and MPS

$$h = 0.68, \Omega_b = 0.048, \Omega_{DM} = 0.23, \Omega_k = 0.001$$

CMB



MPS



Cosmology from quantum fluctuations of space-time (CMaDE model)

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- Conclusions: Quantum fluctuations in space-time seem to be able to explain very well the accelerating expansion of the universe, without the need for a cosmological constant, exotic extra matter, or modifications to Einstein's equations.

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by Tonatiuh Matos (tonatiuh.matos@cinvestav.mx)



$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \frac{2\pi^2}{\lambda^2}g_{\mu\nu} = \kappa^2 T_{\mu\nu}$$

λ = The Compton wavelength of the graviton

Friedman equations

$$\mathcal{M} = \frac{2\pi^2}{\lambda^2}$$

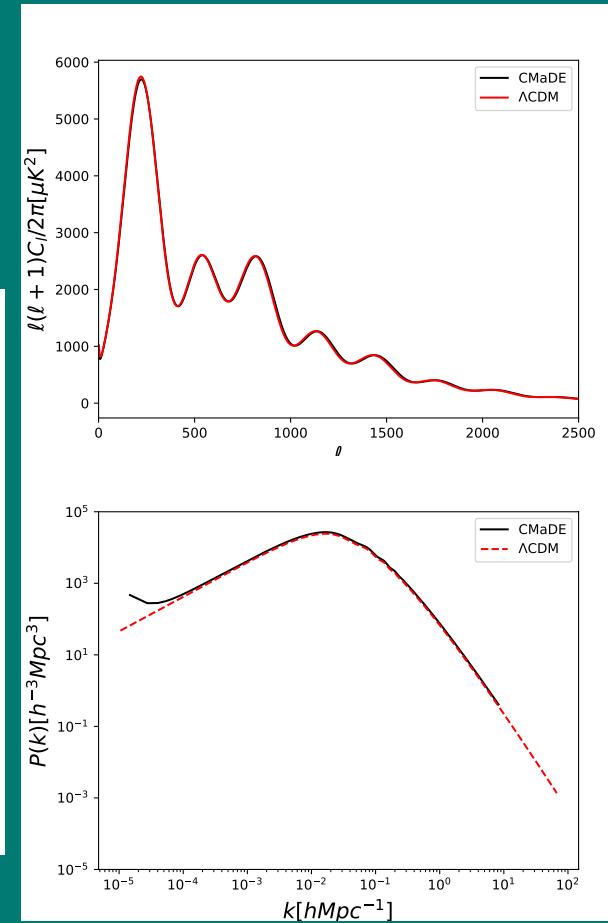
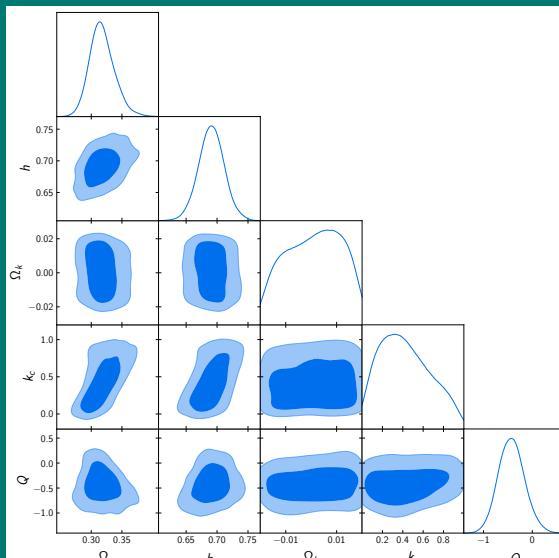
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$$\text{CMB, MPS } h = 0.68, \quad \Omega_b = 0.048 \quad \Omega_{DM} = 0.23, \quad \Omega_k = 0.001$$