

New Cosmological Analysis of the eBOSS Lyman- α Forest

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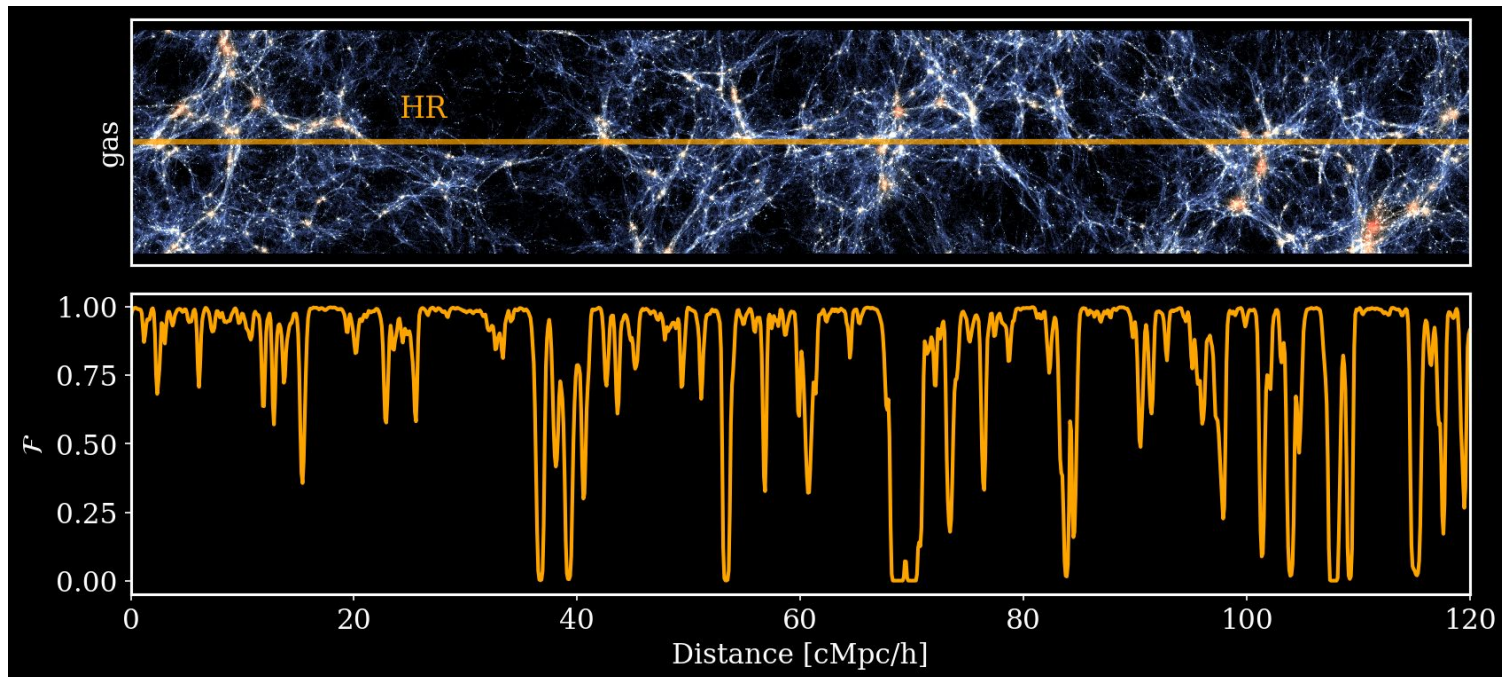
Mahdi Qezlou

arxiv:2306.05471



What is the Lyman- α forest?

Absorption from neutral hydrogen in quasar
Traces dark matter at $z=2-5$



eBOSS SDSS DR14

1D Flux power along line of sight to quasar

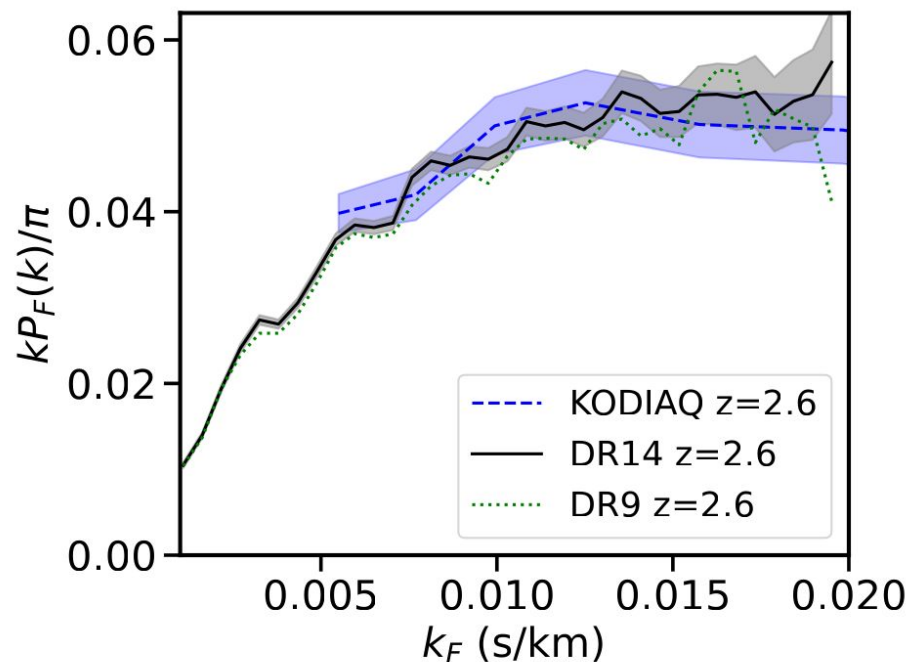
$v = H(z) (a r)$

At $z=3$:

$k = 0.001 \text{ s/km} \sim 0.1 \text{ h/Mpc}$

$k = 0.02 \text{ s/km} \sim 2 \text{ h/Mpc}$

Small non-linear scales!



PRIYA Cosmology Suite

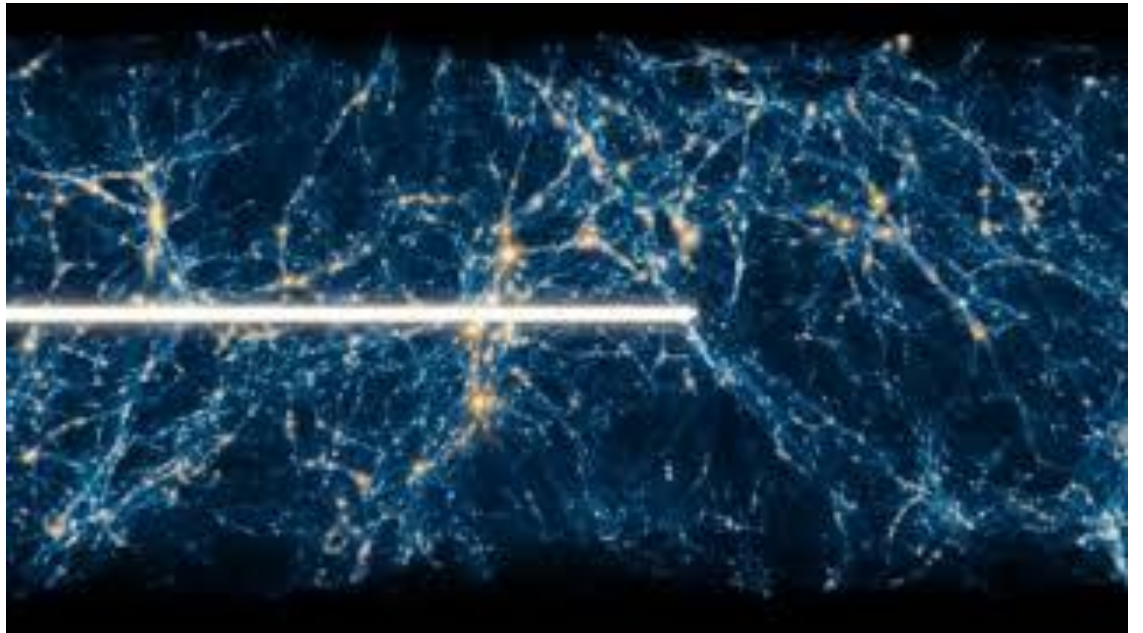
Simulations to model growth and gas pressure

3 High fidelity: 3072^3

48 Low fidelity: 1536^3

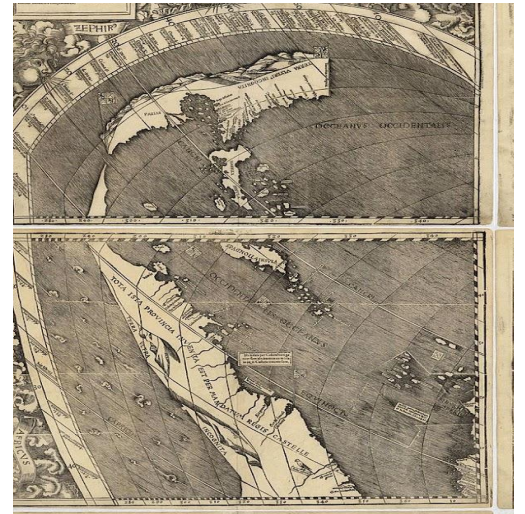
120 Mpc/h box

ASTRID galaxy model



Simulation Interpolation

- For inference need simulation output for all cosmologies
- Gaussian Process interpolation using ~ 50 simulations
- Interpolation needs 50 large simulations

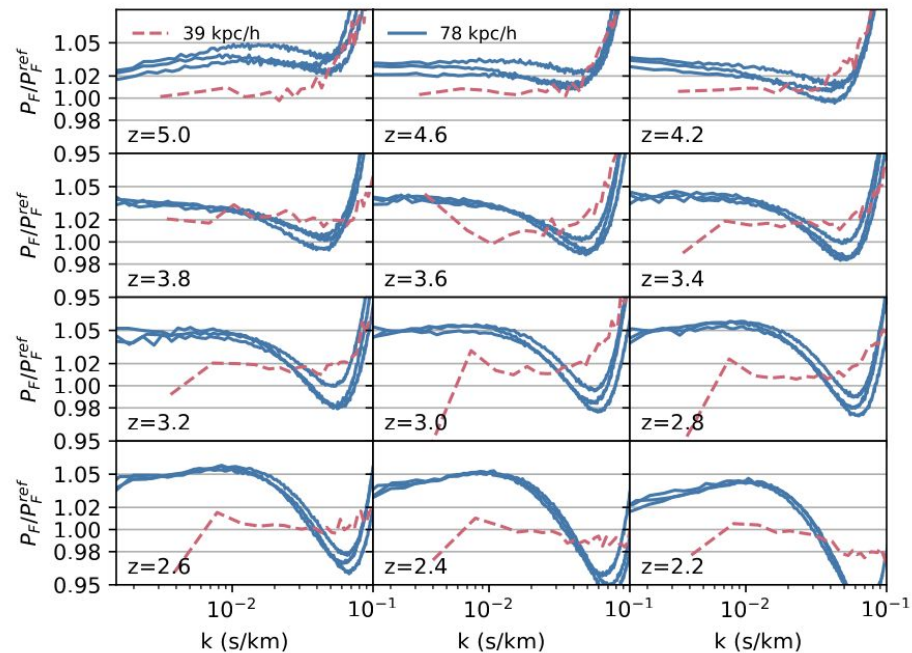


Multi-Fidelity Emulation



Combine simulations at **different resolutions.**

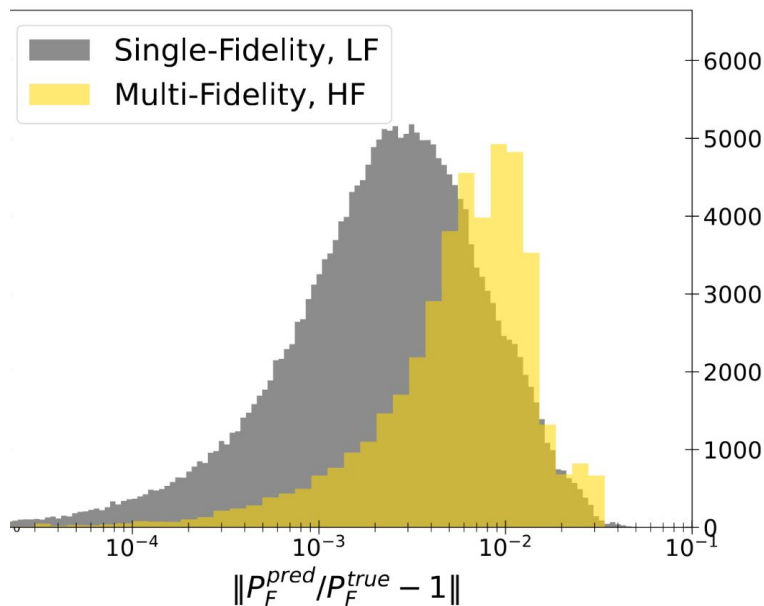
- Low resolution for cosmology
- Correct with high resolution.
- **Cosmology-dependent** correction function.



Leave-one-out Validation



Emulation is 1% accurate! Multi-fidelity leave-one-out is missing $\frac{1}{3}$ simulations



Lyman-alpha Forest Multi-Fidelity



Parameters:

Power spectrum:
$$P(k) = A_P \left(\frac{k}{0.78h/\text{Mpc}} \right)^{n_P-1}$$

Hubble parameter is included, but badly measured

$\Omega_M h^2$ - Growth rate

Helium & hydrogen reionization

Lyman-alpha Forest Multi-Fidelity



Circle: HF. Cross: LF.

Parameters:

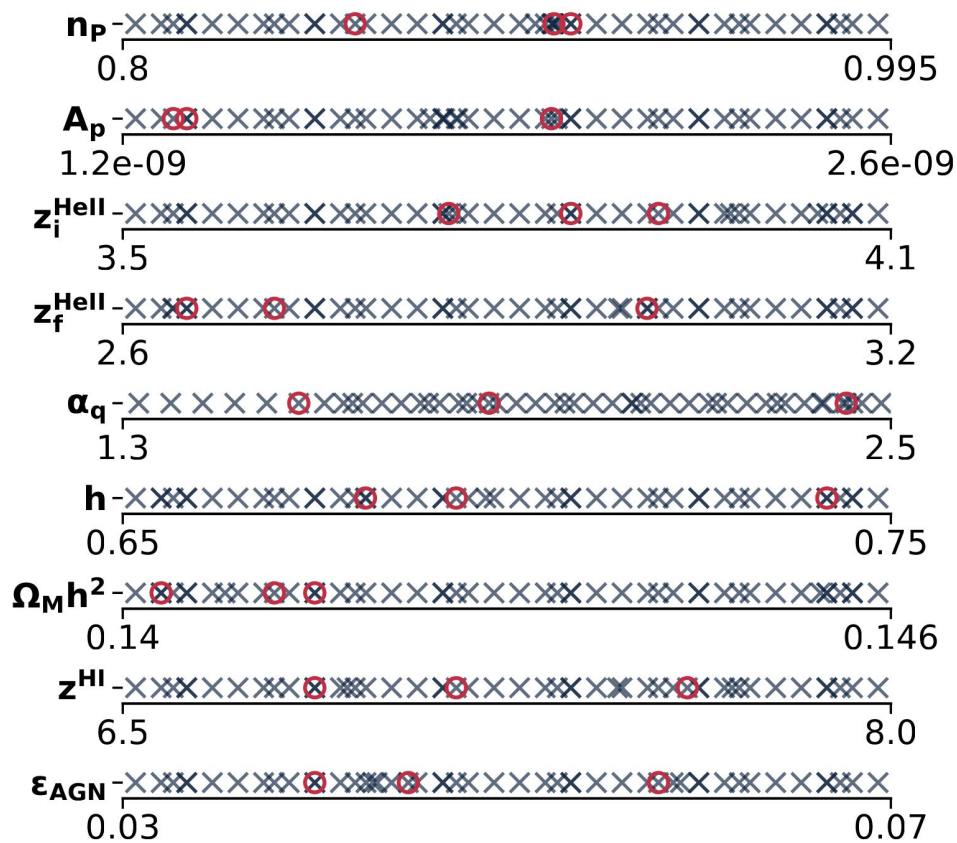
A_p n_p - cosmology

z_i z_f α_q - helium reionization

z_{HI} - hydrogen reionization

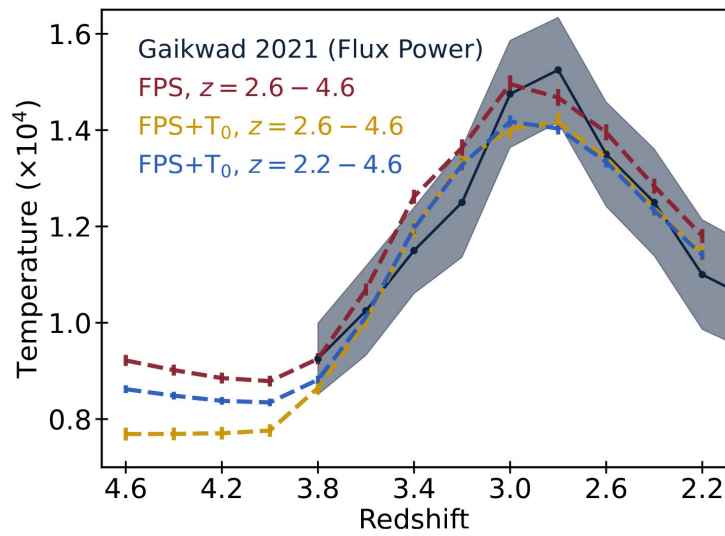
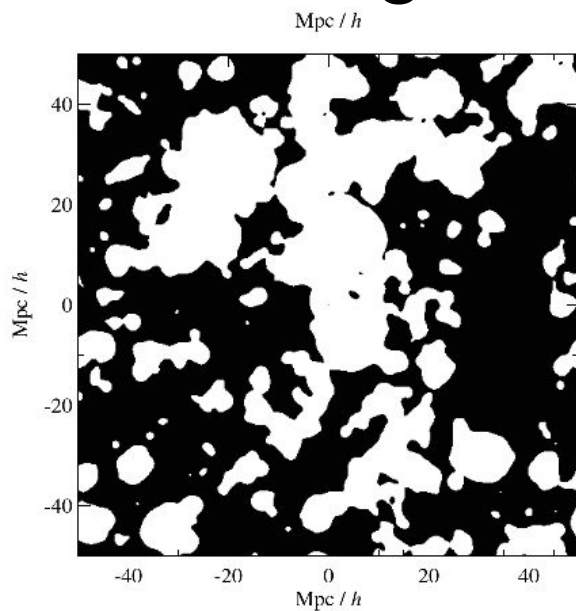
$\Omega_M h^2$ - Growth rate

ϵ_{AGN} - BH feedback



Reionization Models

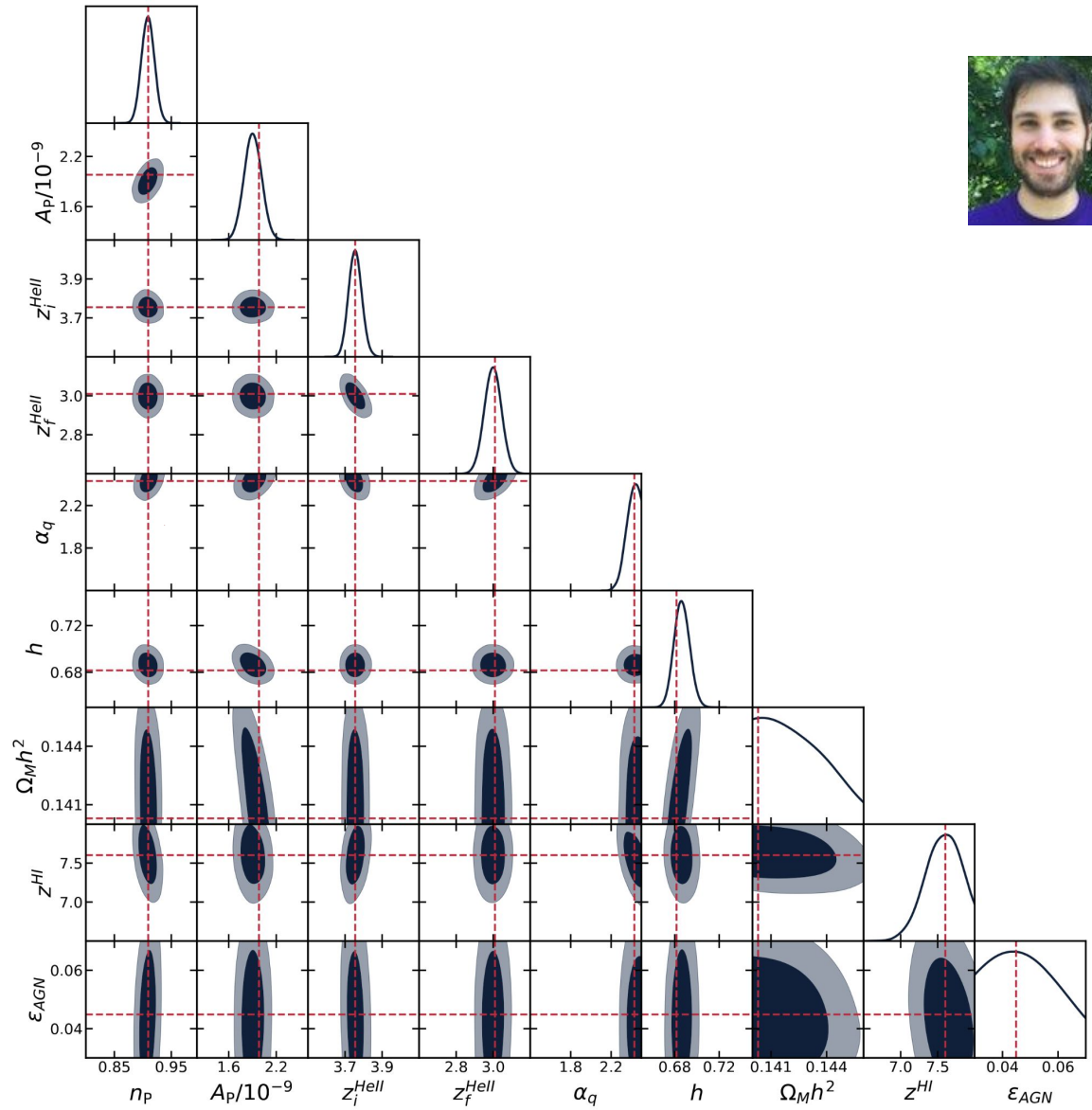
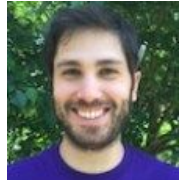
- Patchy hydrogen reionization model
- Patchy helium ~ 30 Mpc bubbles $z \sim 3.8 - 2.8$
- Match gas temperature history



Likelihood function

Posterior constraints with simulated data

- Included:
- Metals
 - DLAs
 - Temperature data



eBOSS SDSS DR14

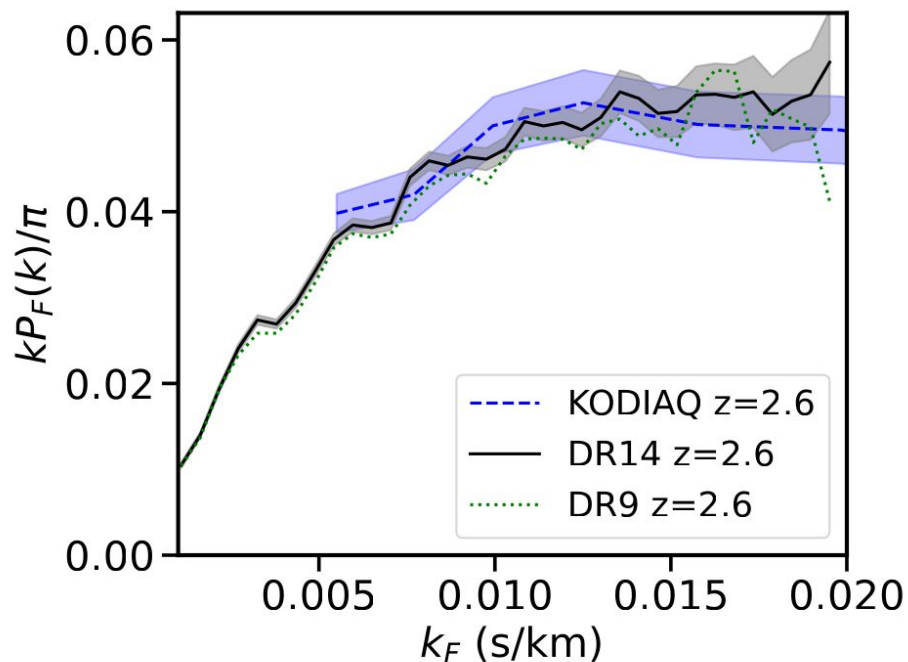
Prefers spectral slope

$$n_s = 0.954 \pm 0.006$$

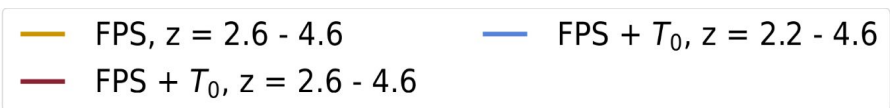
Planck has:

$$n_s = 0.965 \pm 0.004$$

Slight 2σ tension: why?



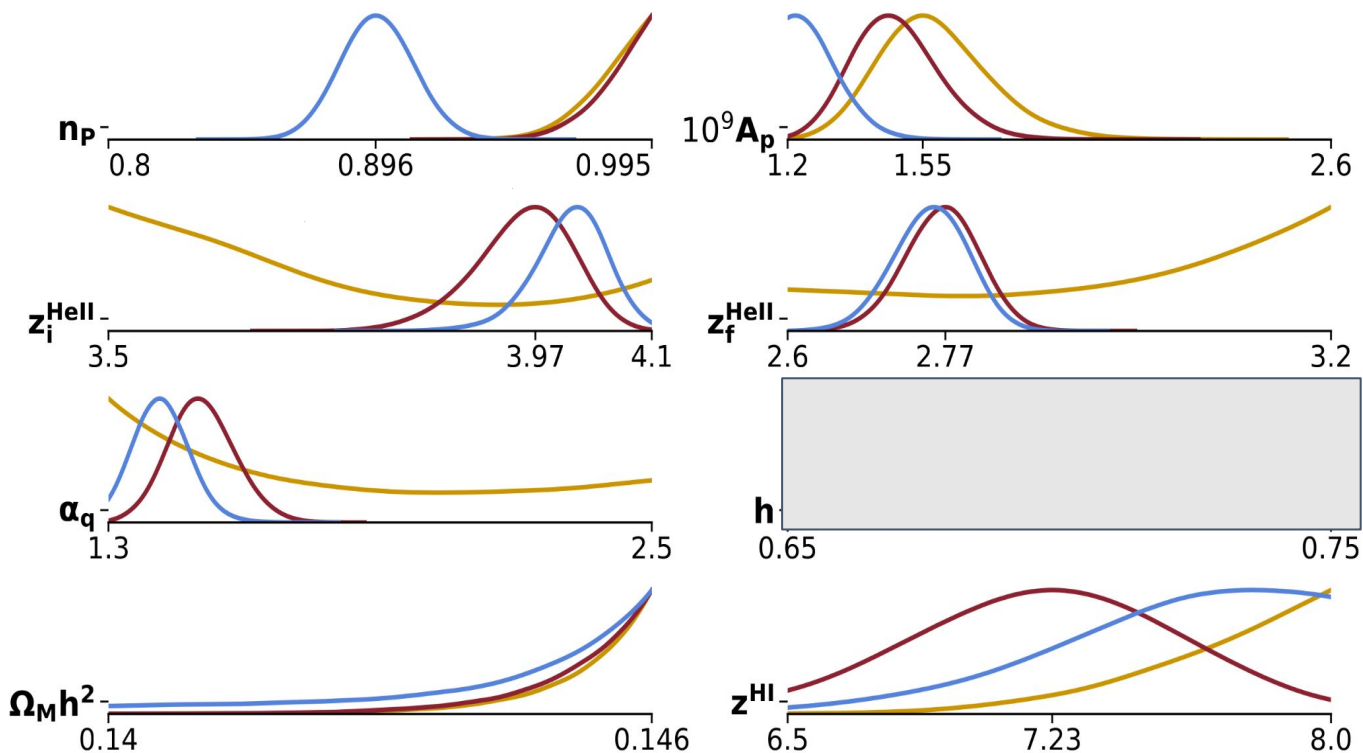
Preliminary DR14



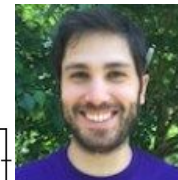
DR14 data:

$n_p \sim 0.9!$ Too low!

Driven by $z = 2.2, 2.4$



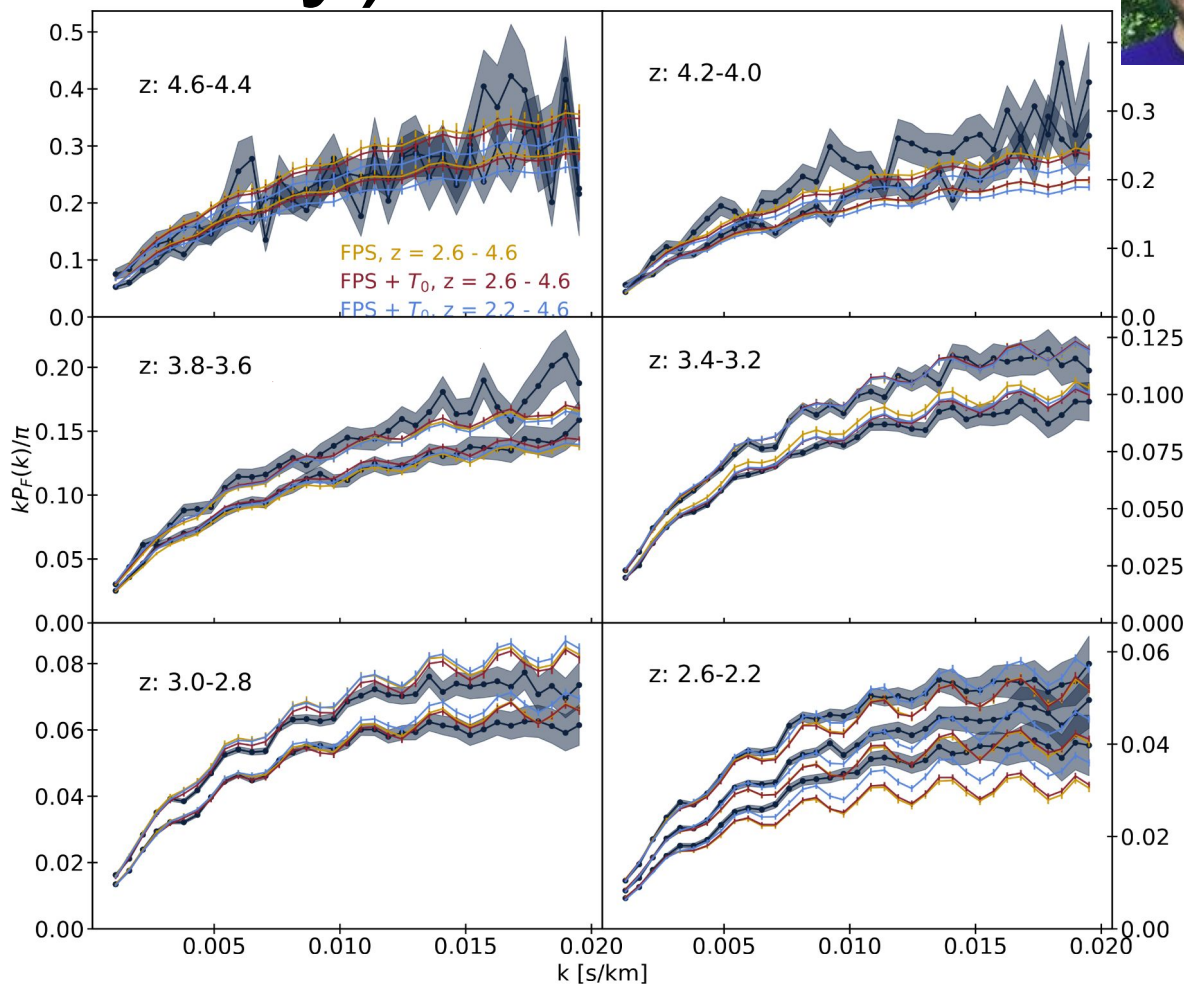
(Preliminary) DR14



Best-fit power:

$z = 2.2$ & 2.4 are
discrepant with
model:

AGN feedback?
Systematic?

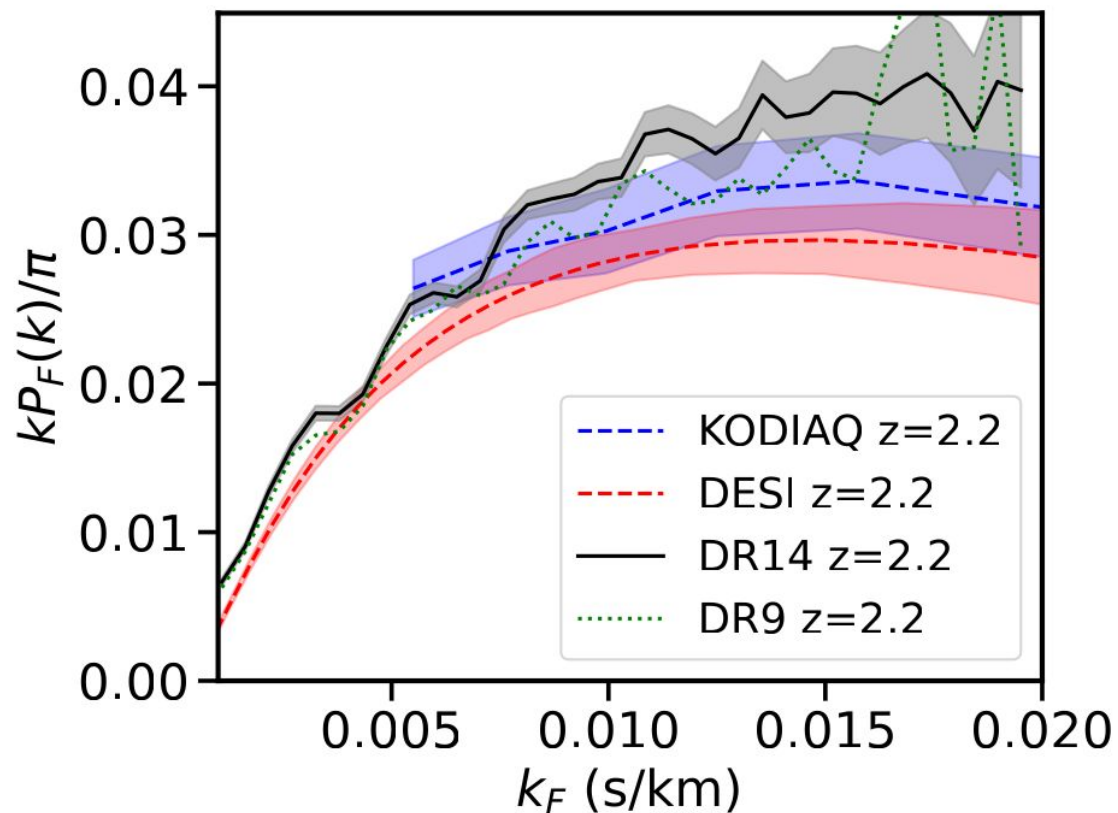


(Preliminary) DR14

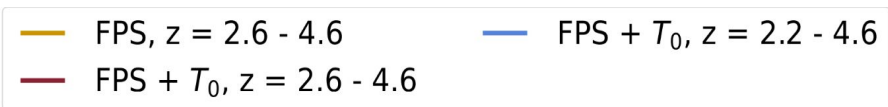


Best-fit power:

Other Lyman alpha data suggests a systematic, but not conclusive



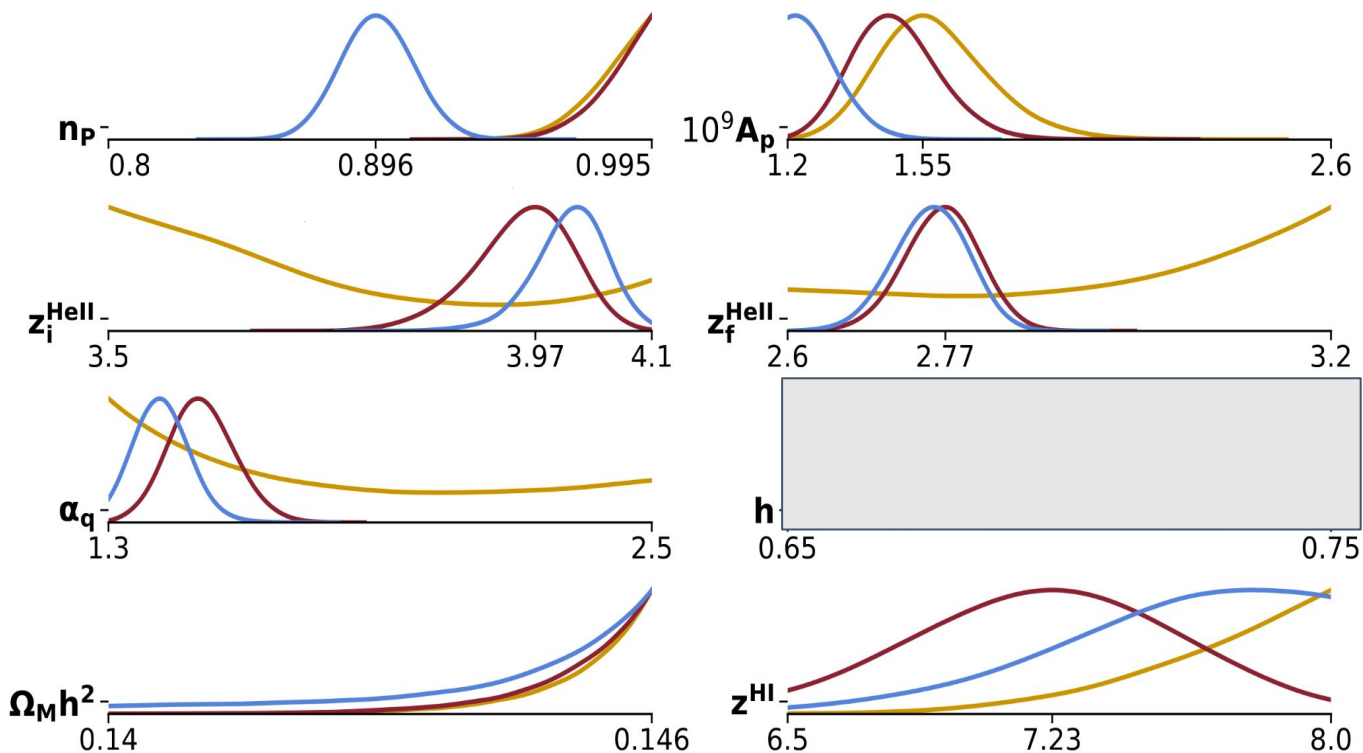
(Preliminary) DR14



DR14 data alone
(not joint chains!)

$n_s \sim 1 - 0.97$

$\sigma_8 \sim 0.8$



Conclusions 2306.05471

Reanalysis of eBOSS
Lyman alpha data

PRIYA suite and
multi-fidelity emulation

Generally weaker
constraints: more
room for theorists!

Until DESI!

