



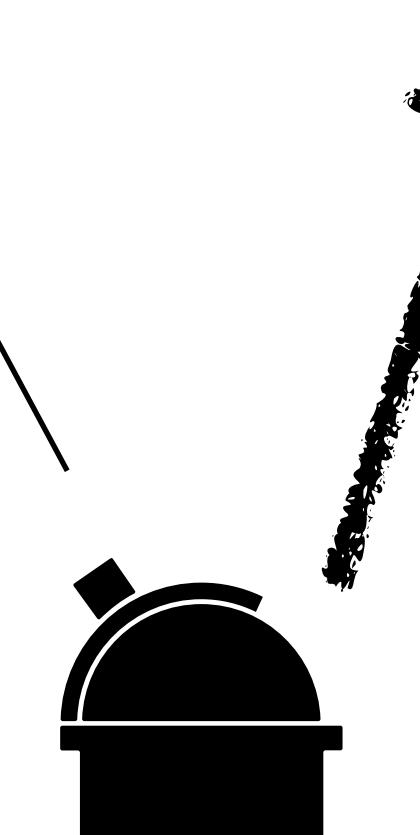
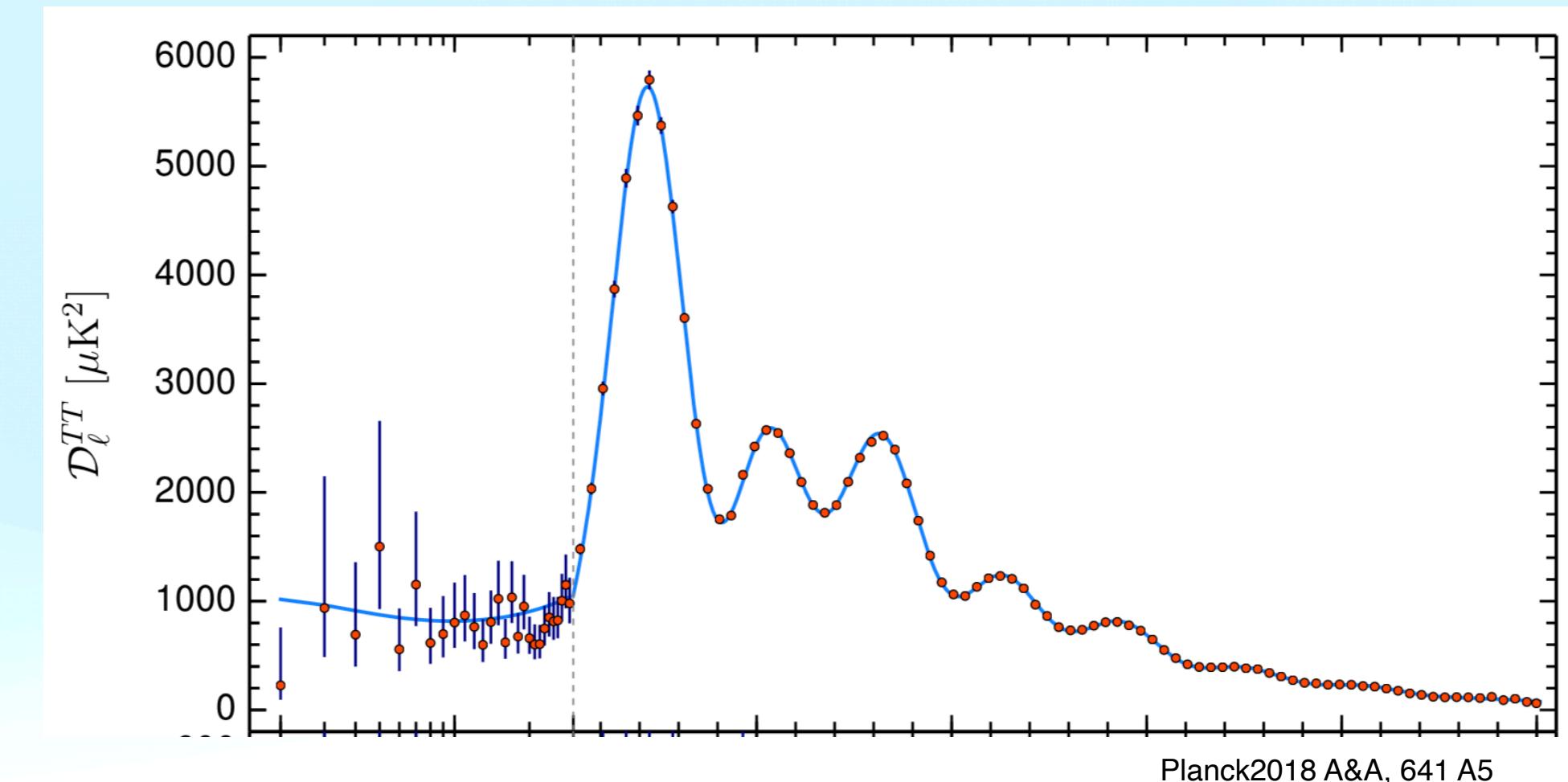
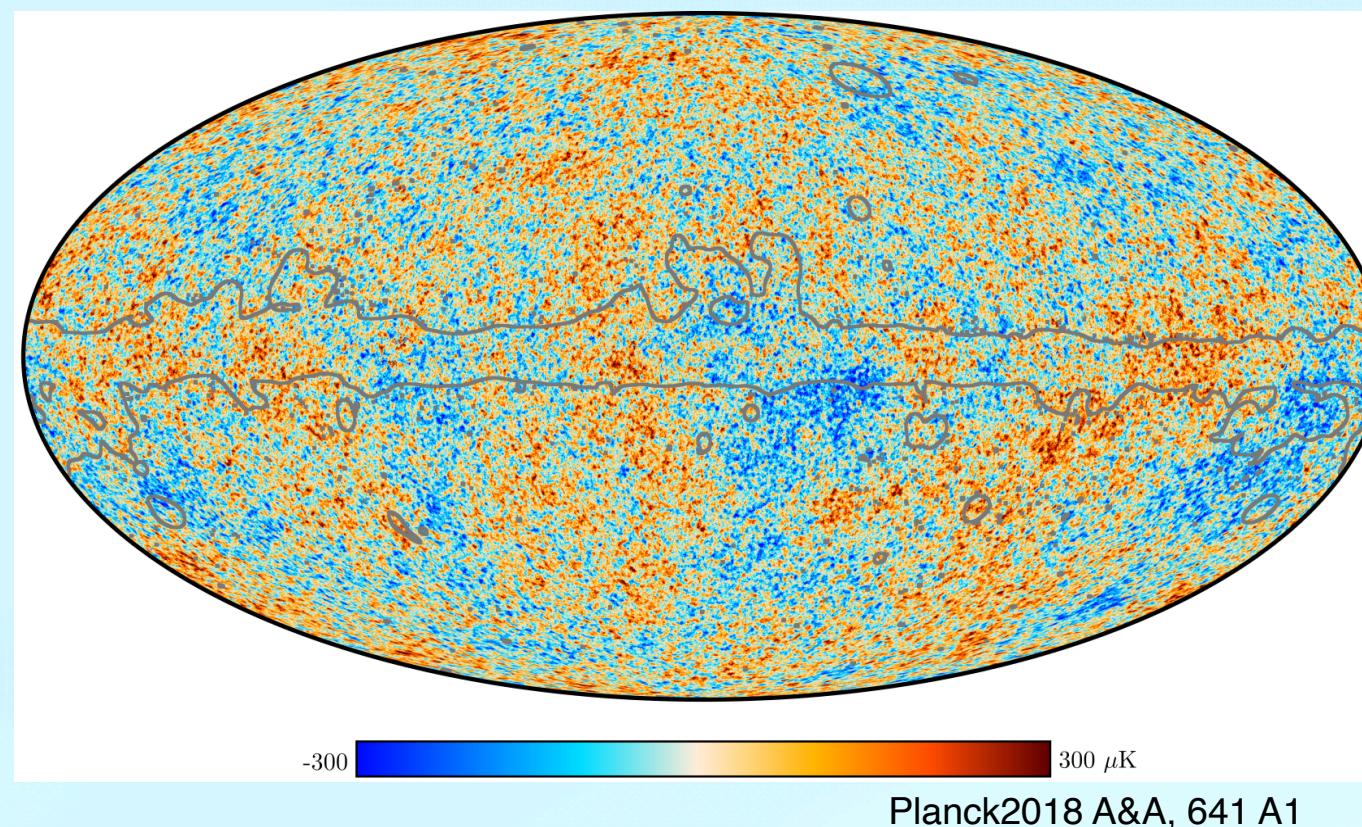
Universiteit
Leiden



Shape of CMB lensing in EDE

arXiv:2305.18873,
Gen Ye, Jun-Qian Jiang, Yun-Song Piao

Cosmological model

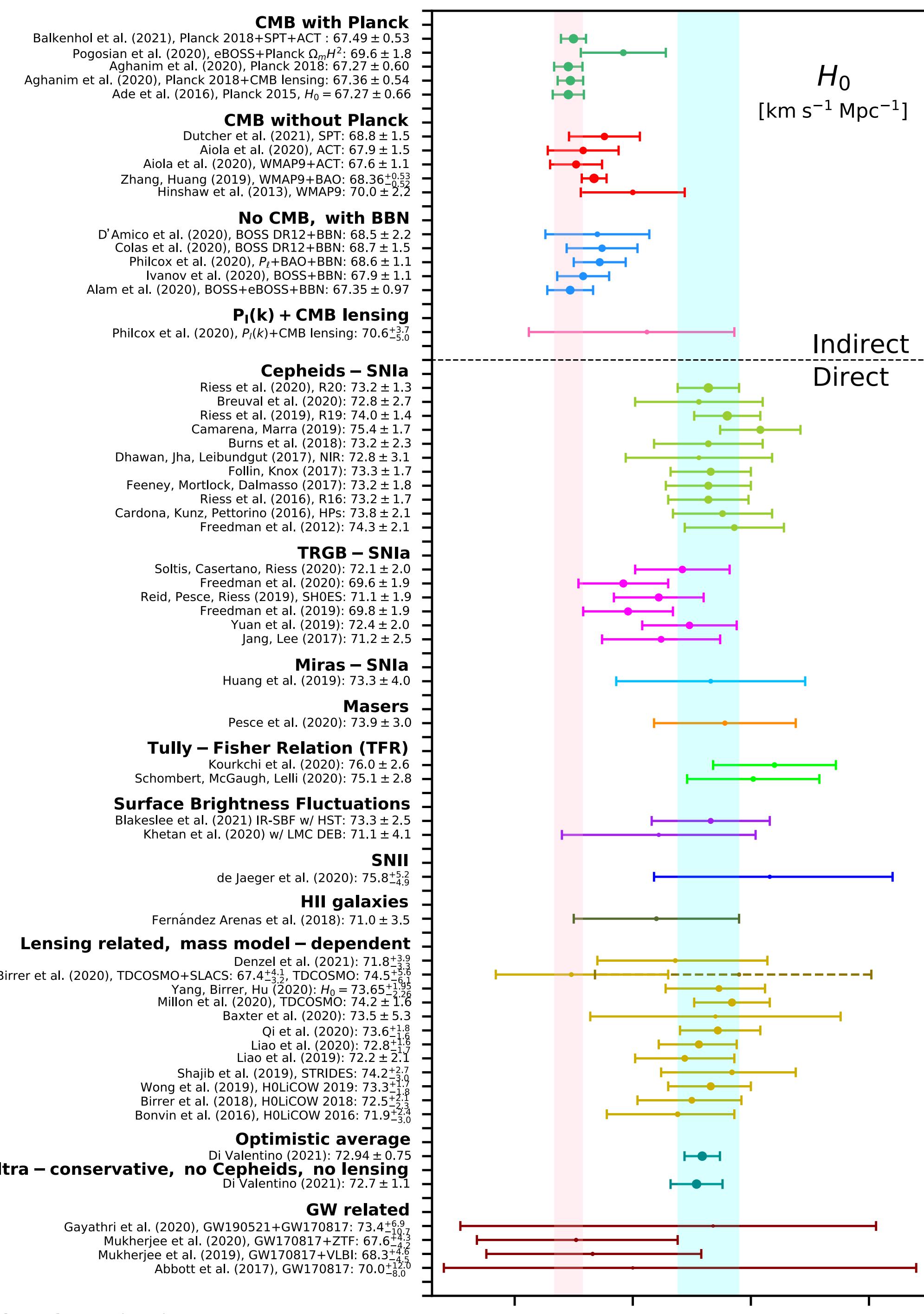


Cosmological Model
e.g. ΛCDM

$\omega_b, \omega_{cdm}, H_0, A_s, n_s, \tau$

Tensions in the model?

Hubble tension



Early Dark Energy (EDE)

$$\theta_s = \frac{r_s}{D_A} \sim \frac{1}{\Delta l}$$

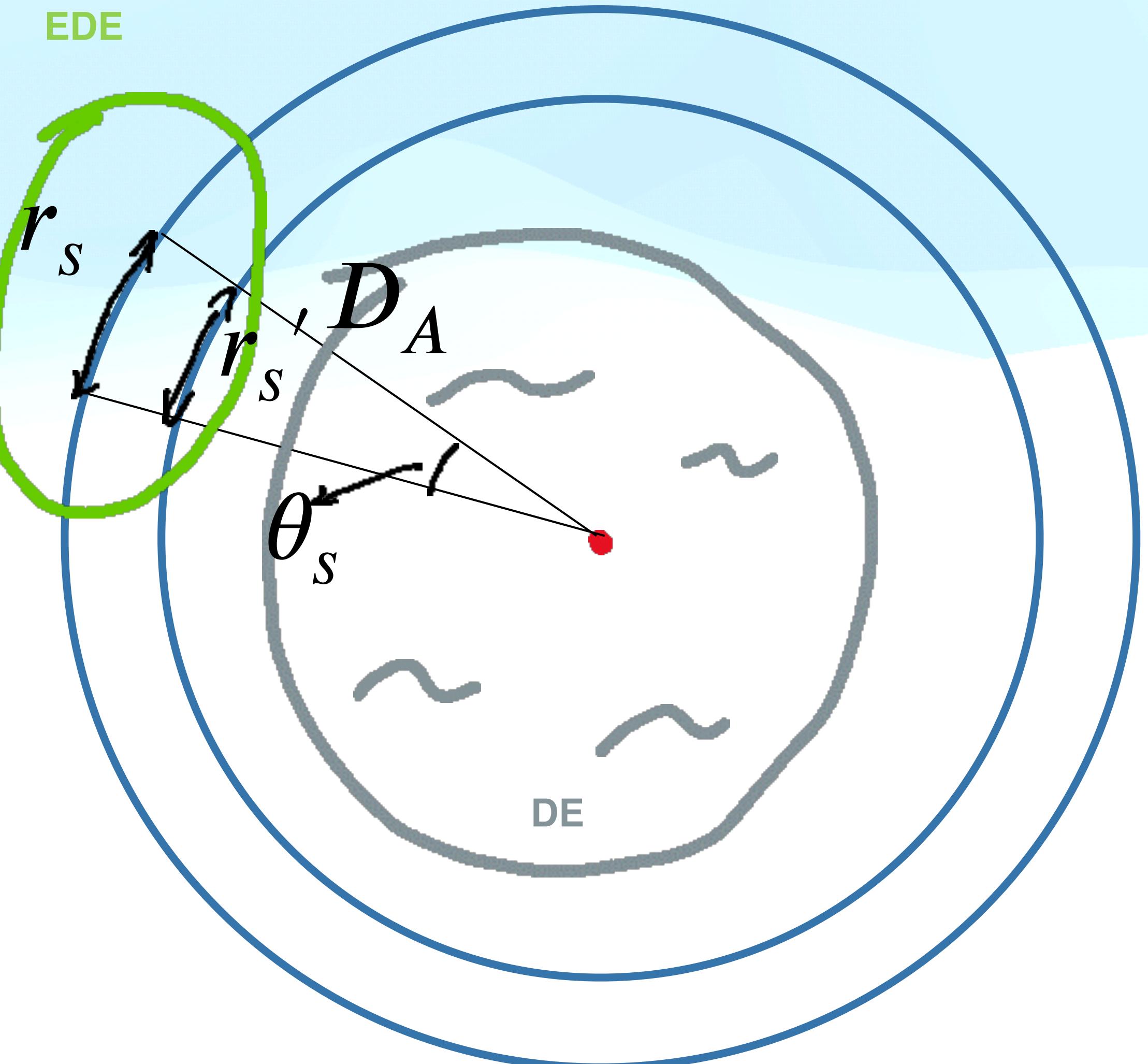
$$r_s = \int_{z_*}^{\infty} \frac{c_s}{H(z)} dz$$

$$D_A = H_0^{-1} \int_0^{z_*} \frac{dz}{E(z)}$$

Tanvi Karwal, Marc Kamionkowski, PhysRevD.94.103523
Vivian Poulin, Tristan L. Smith, Tanvi Karwal, Marc Kamionkowski, PhysRevLett.122.221301

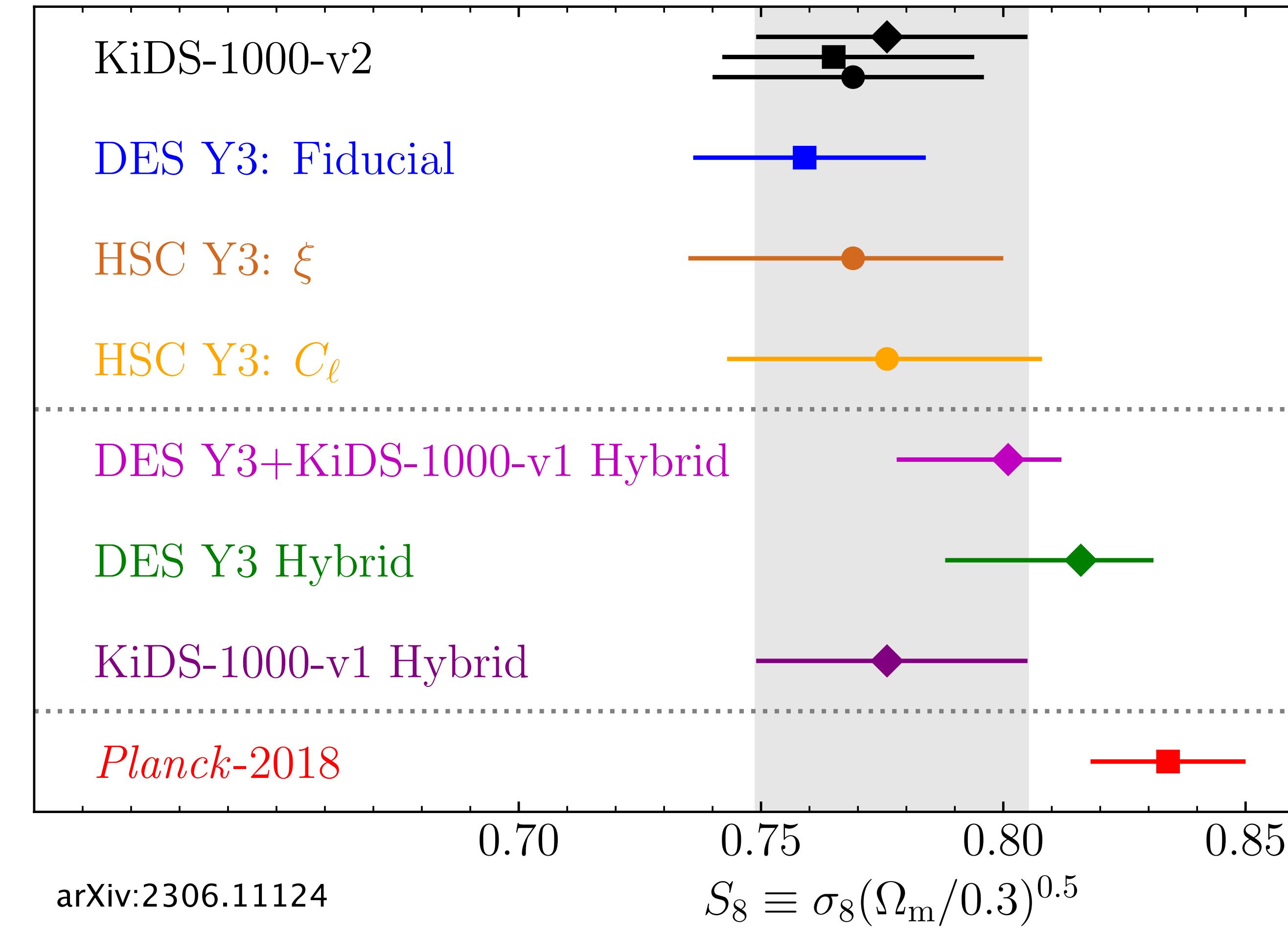
Early Dark Energy

Dark Energy

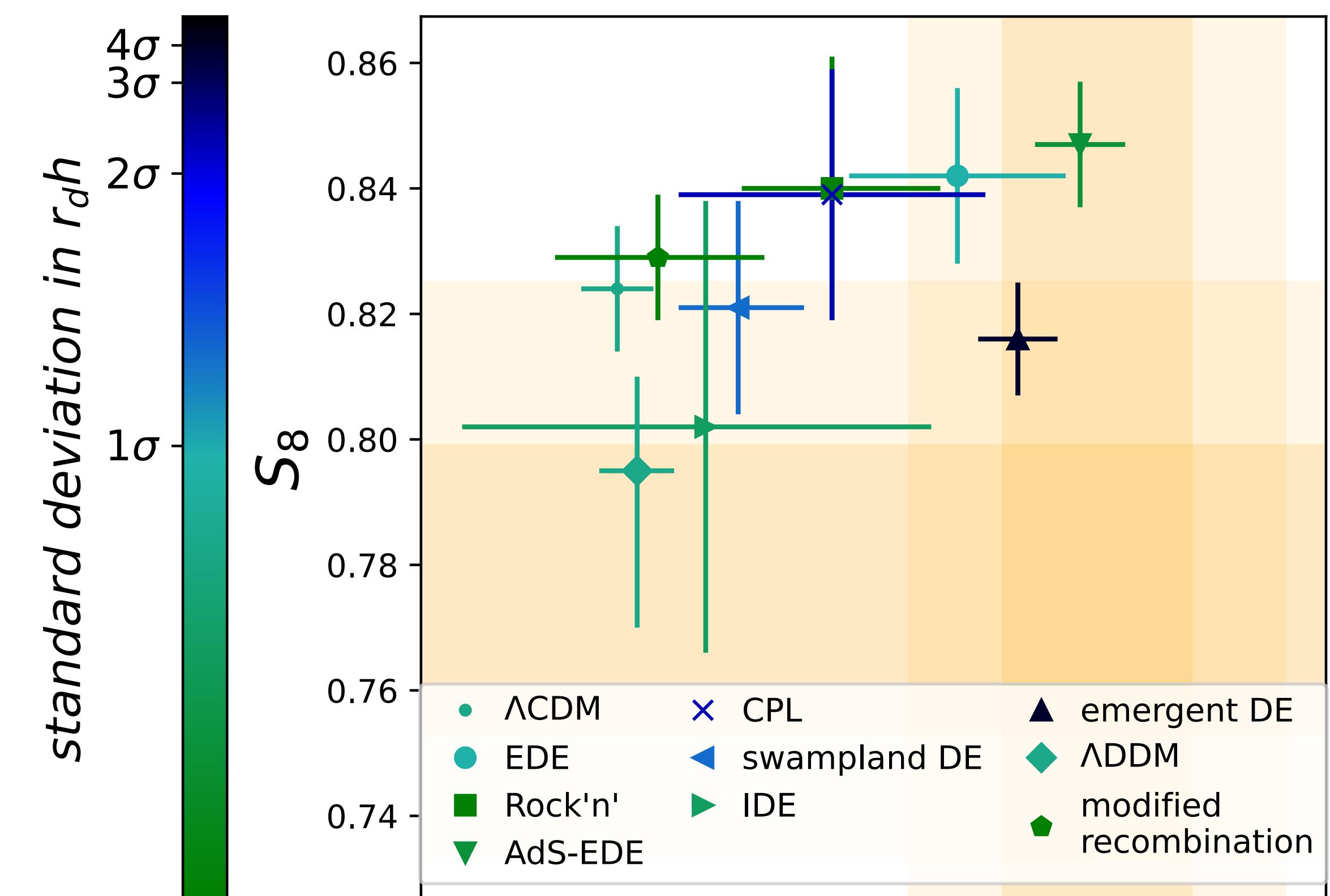
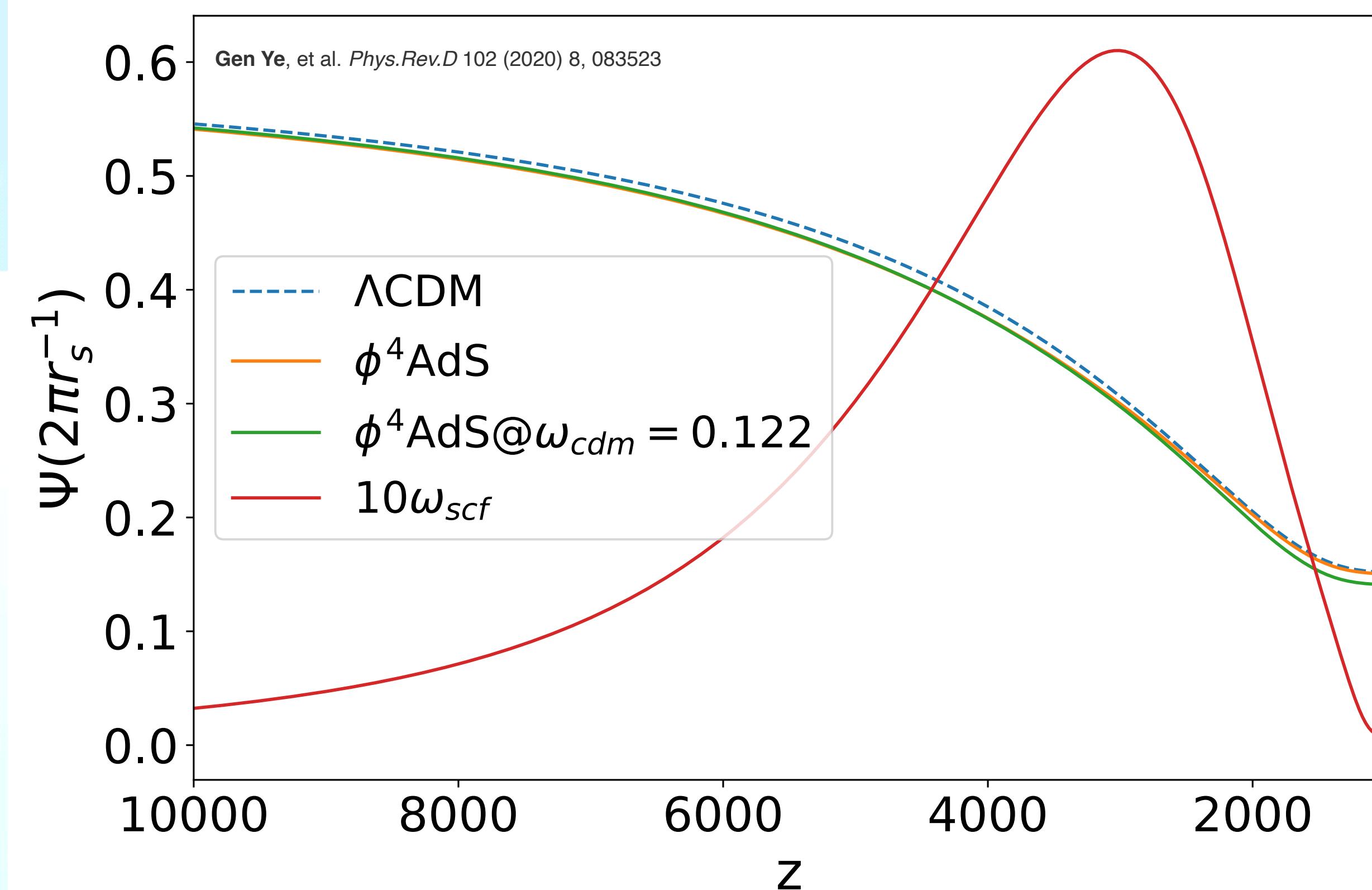


Tensions in the model?

S_8 tension



EDE and S_8



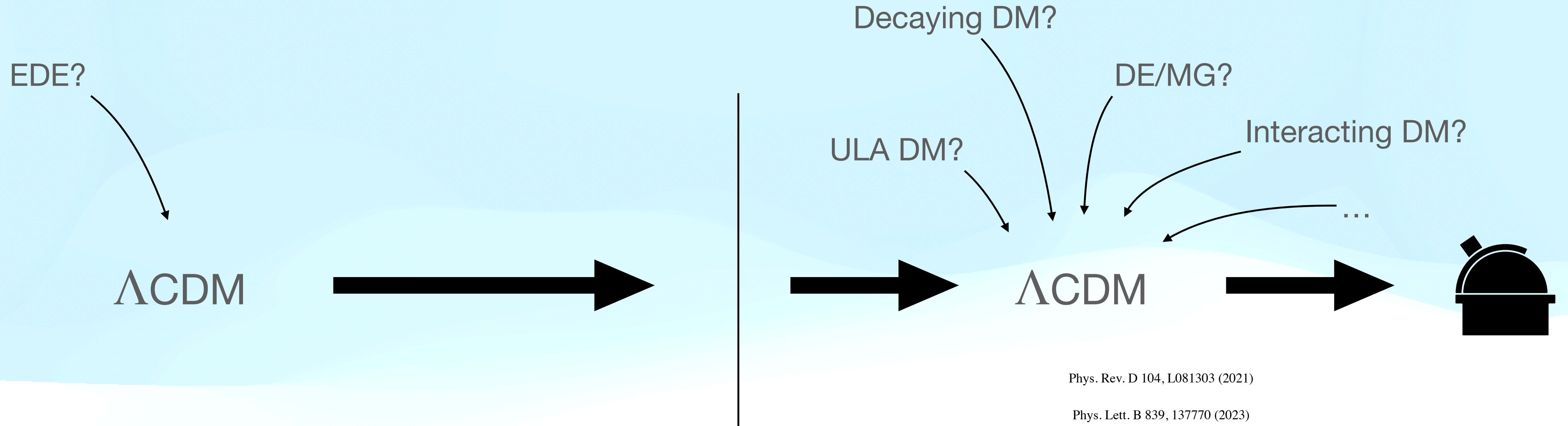
S_8 tension exacerbated by EDE

ω_{cdm} is increased to compensate for EDE's effect
in radiation driving and ISW

Gen Ye, et al. *Phys. Rev. D* 102 (2020) 8, 083523

Astrophys. J. Lett. 904, L17 (2020)

Phys. Rev. D 104, 063524 (2021)



EDE?

Λ CDM

Background compatibility?

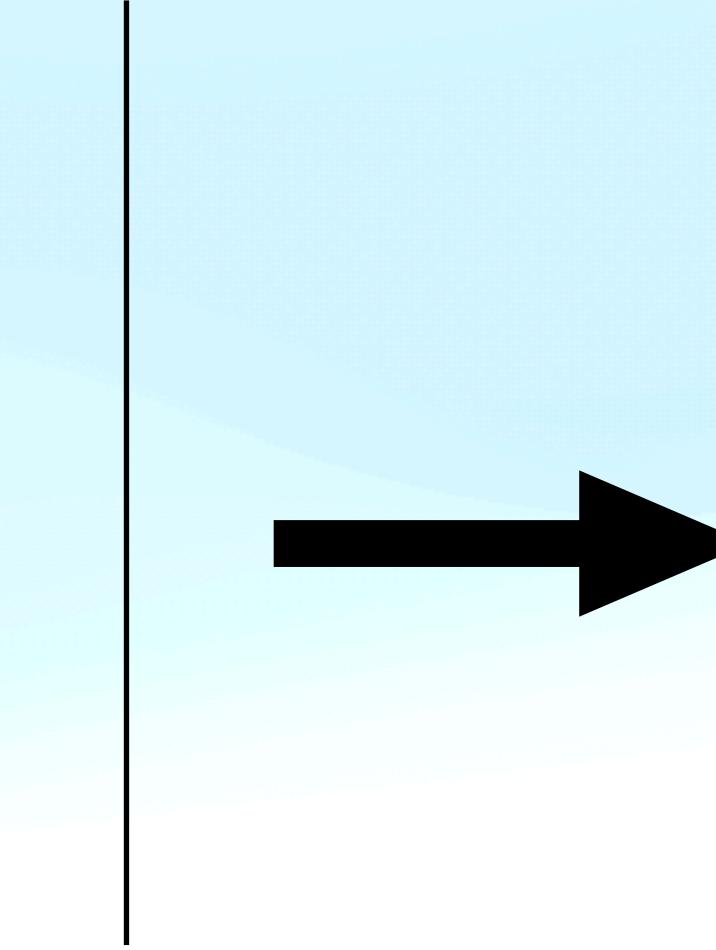
Cosmological parameters

$$D_A = H_0^{-1} \int_0^{z_*} \frac{dz}{E(z)}$$

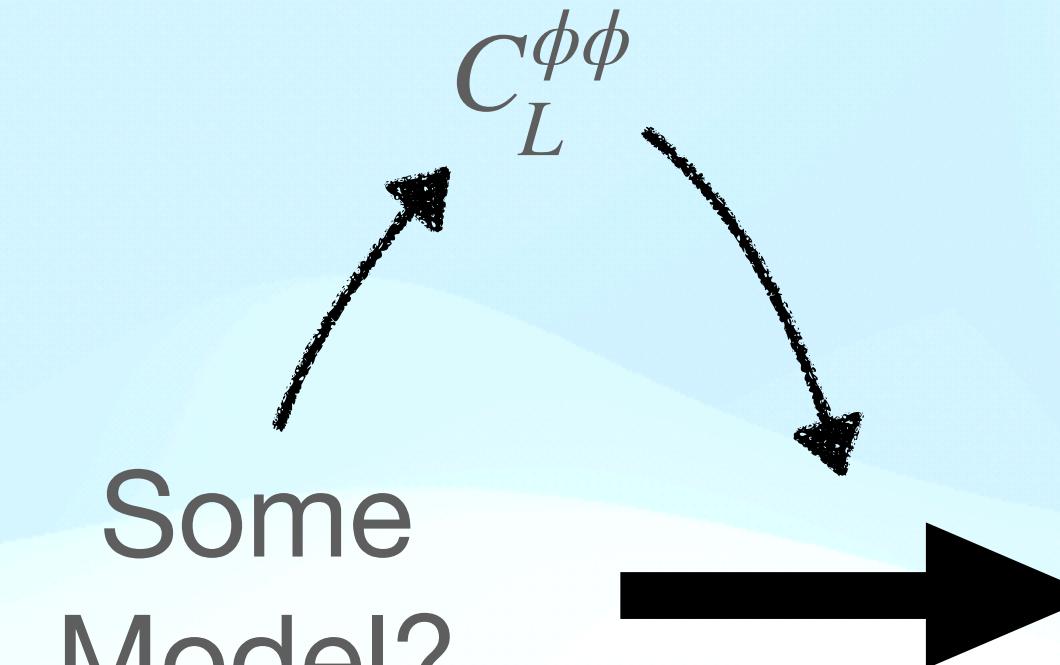
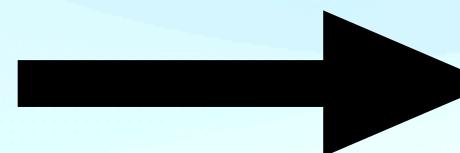
Dark energy EoS $w(z)$

Phys. Lett. B 832, 137244 (2022)

arXiv:2302.07333



Recombination



Some
Model?



Perturbation compatibility?

Reionization

Low ℓ E polarisation

High ℓ optical depth

Gravitational Lensing

This work arXiv:2305.18873

EDE?

Λ CDM

$$C_L^{\phi\phi} = A(L) C_{L,fid}^{\phi\phi}$$

$$A(L) \sim GP[\bar{f}(L), K]$$

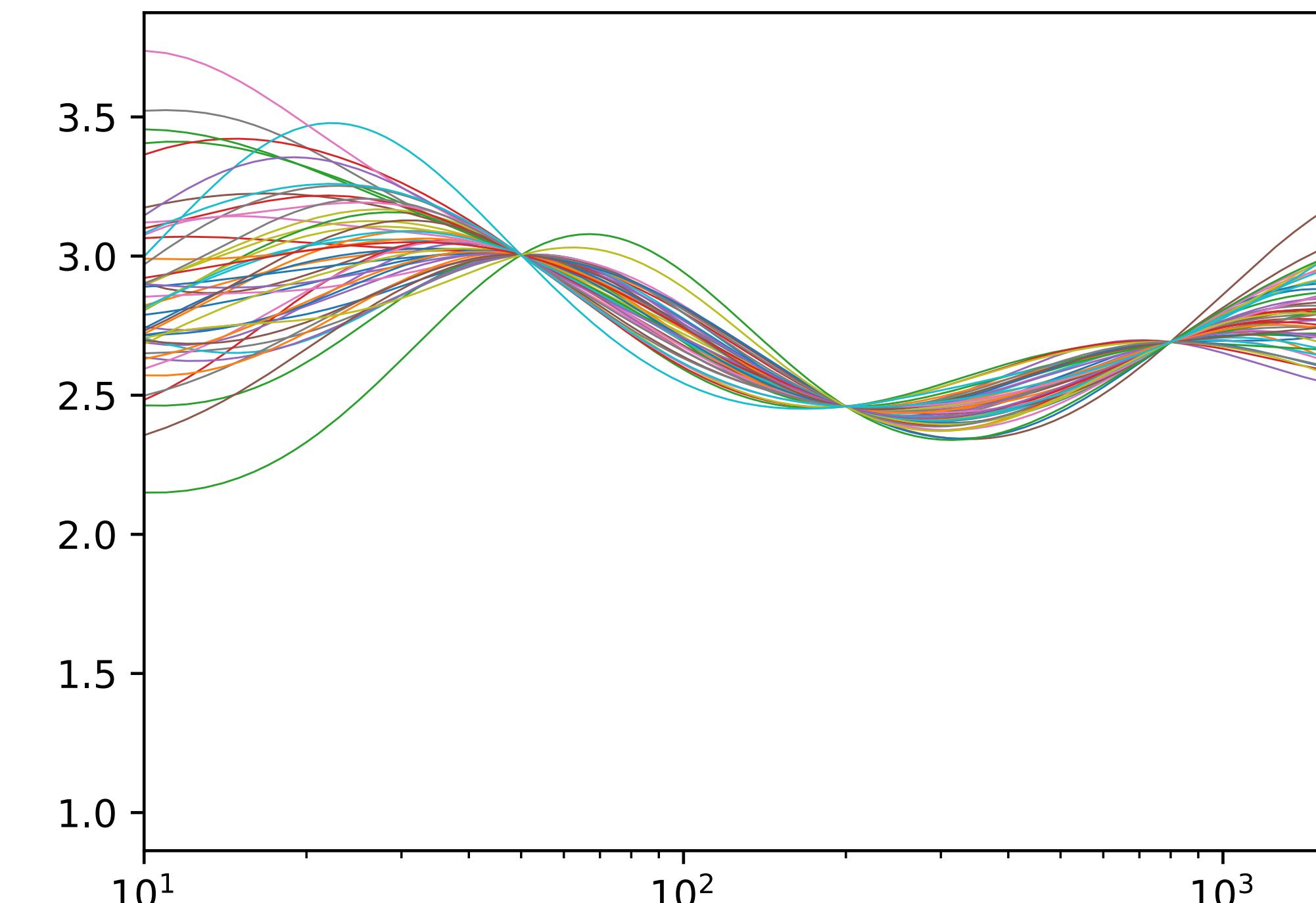
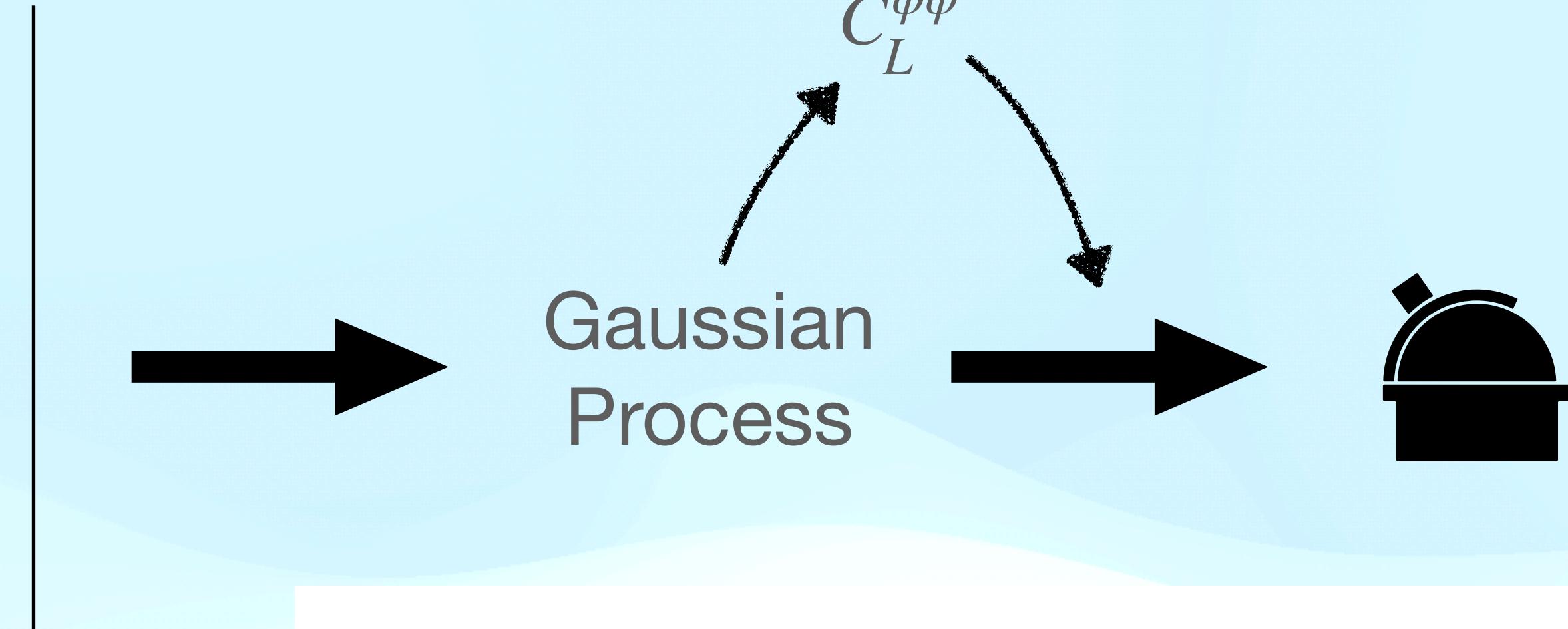
3 Nodes: $\{A_{L=50}, A_{L=200}, A_{L=800}\}$

Previous works: binning+interpolation

Phys. Rev. D, 107 103505, ...

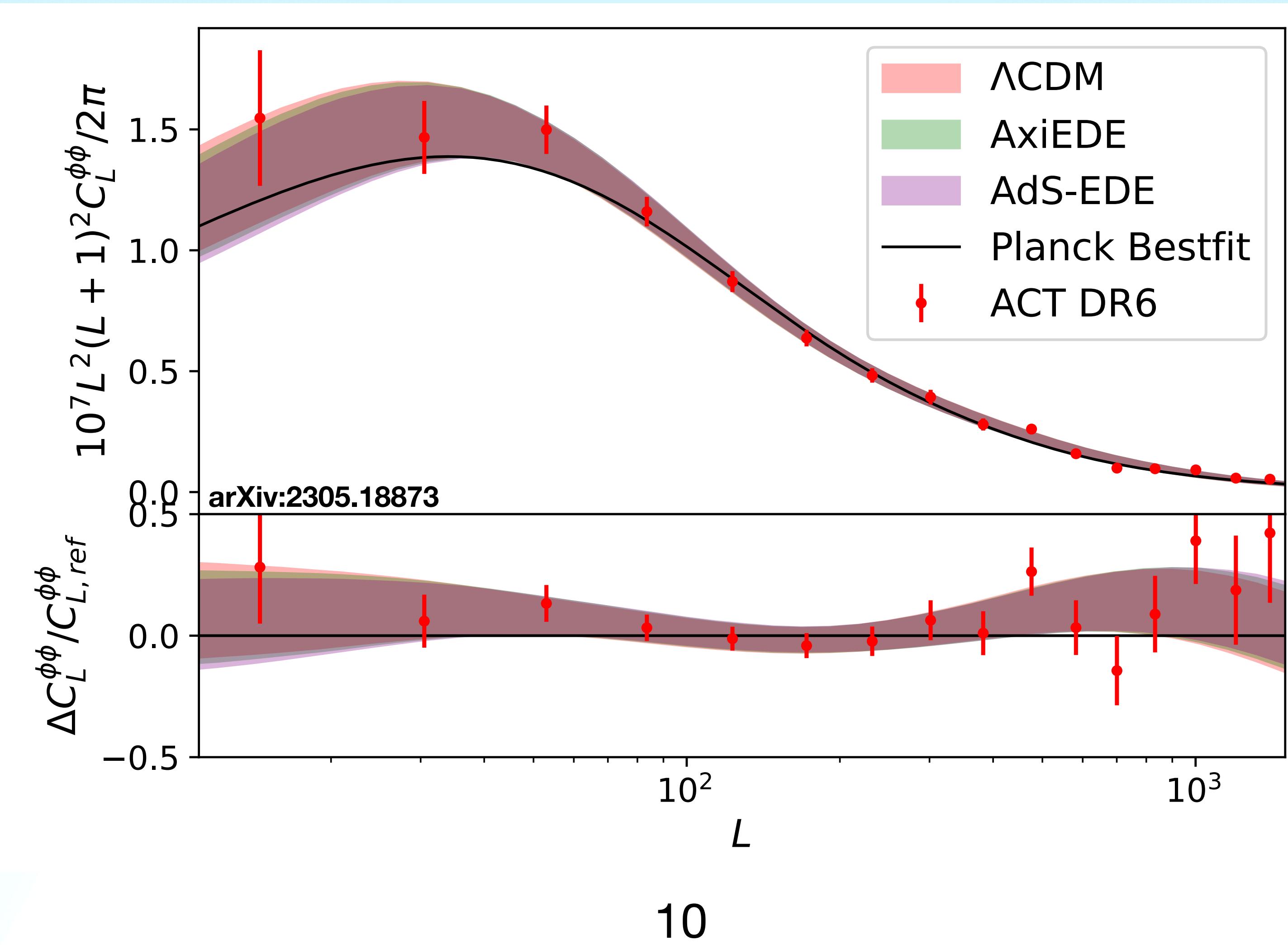
Recombination

9



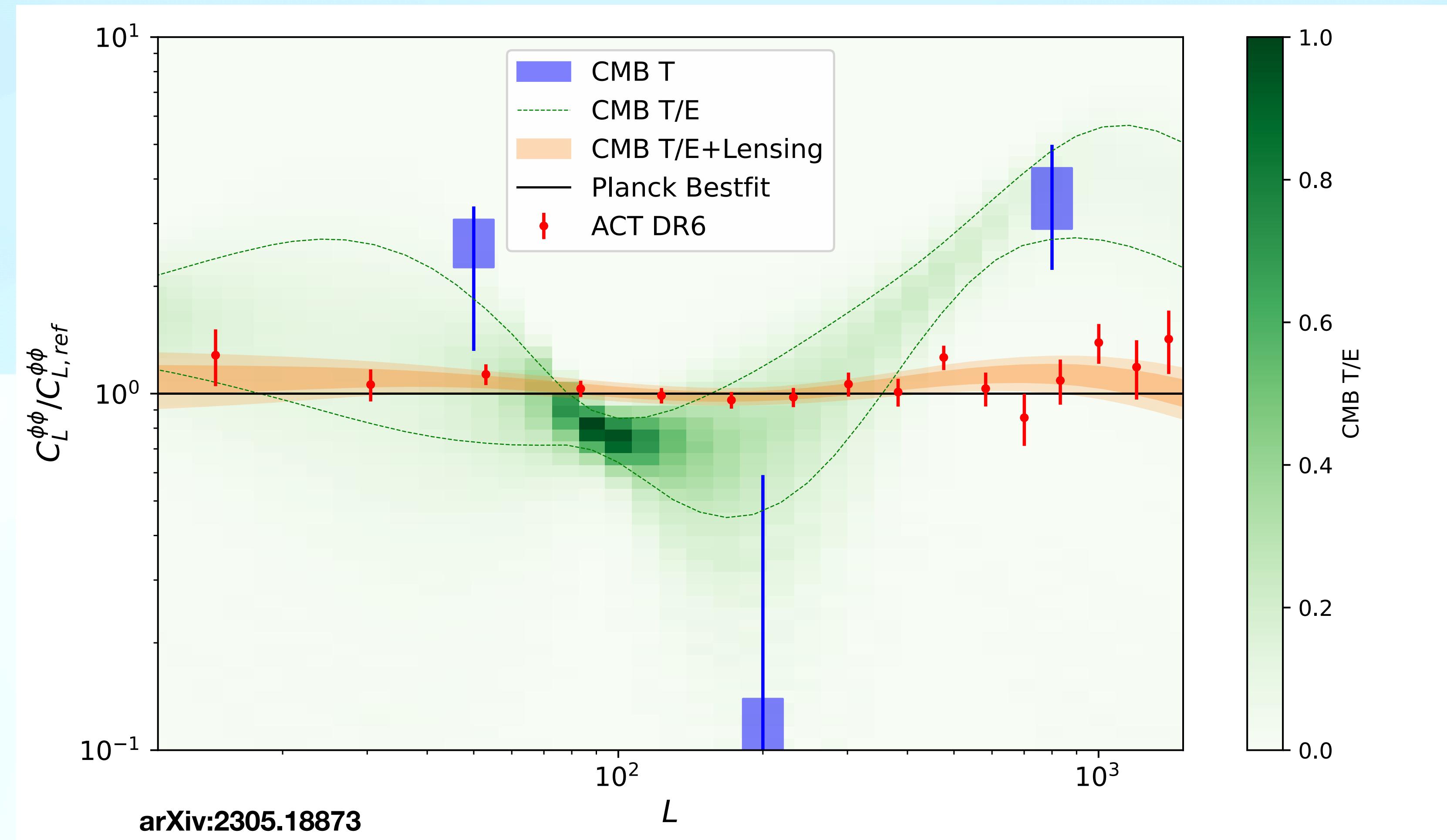
Results

CMB TTTEEE + Lensing reconstruction



Results

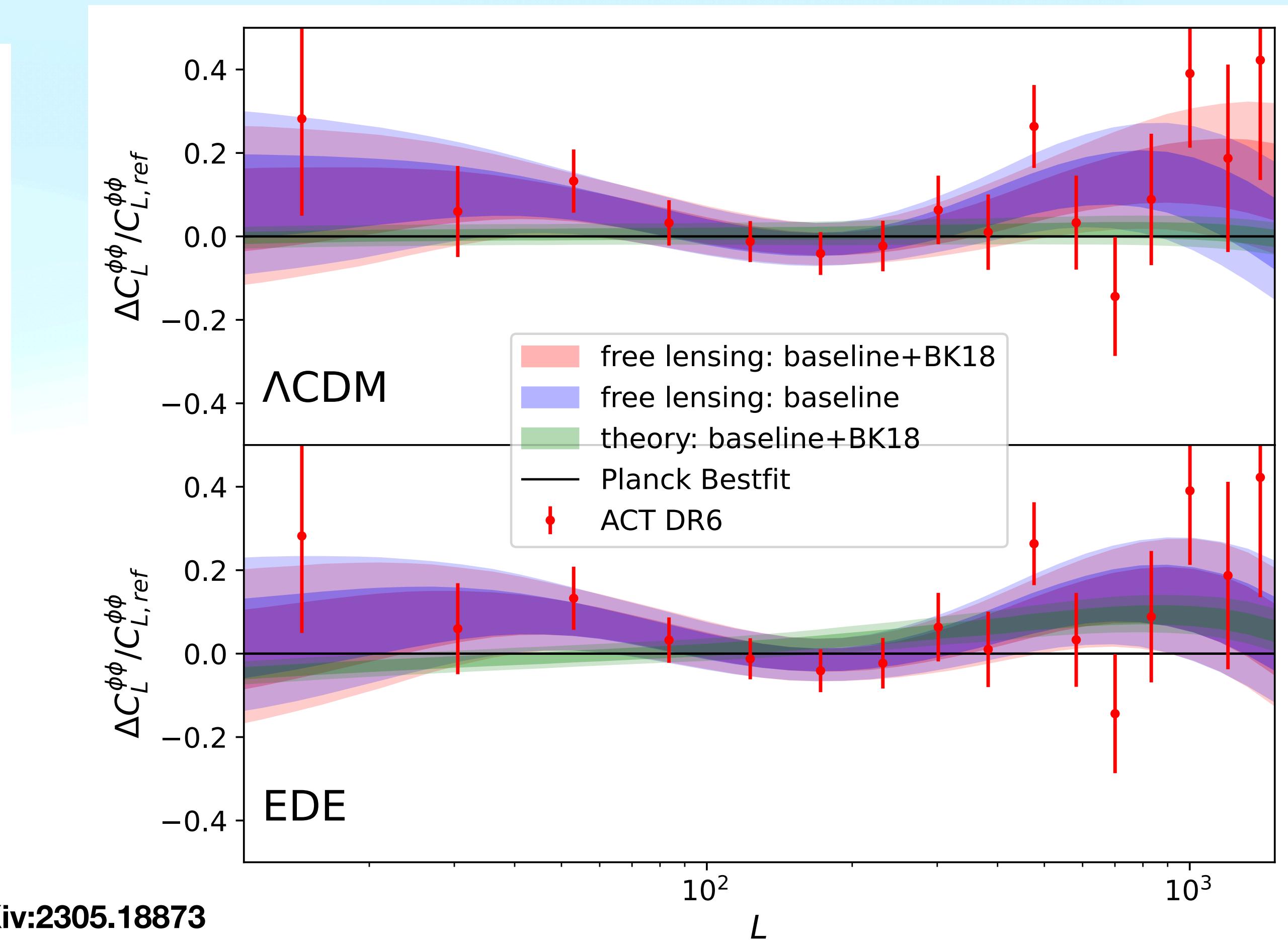
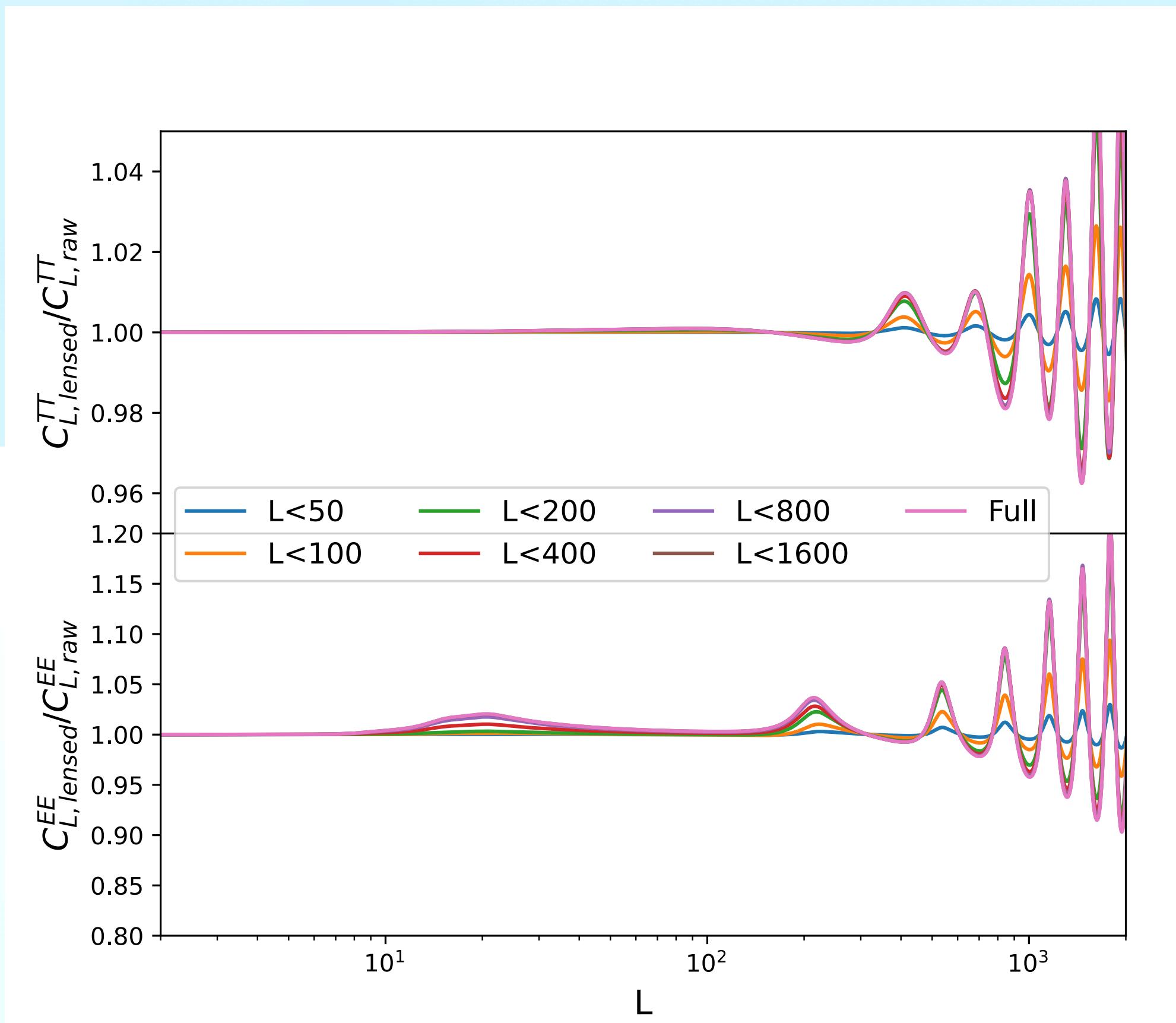
CMB T/E + Lensing reconstruction



arXiv:2305.18873

Results

CMB TTTEEE + CMB B mode + Lensing reconstruction



arXiv:2305.18873

Conclusion

- GP sampling method, fewer nodes required
- Late universe independent constraint on lensing
 - Insensitive to early Universe model (Λ CDM vs EDE)
 - $80 < l < 400$ dominated by lensing reconstruction
 - High and low l pushed up by lensed T/E
 - New constraint from B mode
- Work in progress: crossing with galaxy weak lensing